

Sciothericum Telescopicum;

O R,

A New Contrivance of Adapting a

TELESCOPE

TO AN

Horizontal Dial

FOR OBSERVING

The moment of Time by Day or Night.

Useful in all *Astronomical Observations*, and for
Regulating and Adjusting

Curious *Pendulum-Watches* and other *Time-Keepers*,
With proper *Tables* Requisite thereto.

By *William Molyneux* Esq; Fellow of the *Royal Society*,
and of that in *Dublin*.

D U B L I N,

Printed by *Andrew Crook* and *Samuel Helsham*, for
William Norman, *Samuel Helsham* and *Eliphal Dobson*
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Advertisement to the Book- Binders.

THat the last Leaf in the Sheet *L* of the
Tables is to be Cut off, and turned, and
the side marked with the Star is to be placed
next to the sixth Page.

TO HIS
EXCELLENCY

Henry Earl of Clarendon,
LORD Lieutenant General,

AND
GENERAL GOVERNOUR
OF

His MAJESTIES Kingdom
OF
IRELAND.

THe great Honour Your Excellency was pleased
to shew Our Society in accepting Our Congratu-
lation at Your happy Arrival and Settlement in the Go-
vernment of this Kingdom, and in giving us Your In-
couragement in prosecuting our Philosophical Designs,
does

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does necessarily require from us all possible Returns of Gratitude. And though what I now Present Your Excellency may seem unfit to begin withal, yet I could not omit this opportunity of expressing to Your Lordship the deep sense I have of Your Favours to us, and to Philosophy. And certainly, my Lord, True Philosophy does highly deserve Encouragement from the Wise, the Great, and Powerful; the most Learned Universities have at all times made Natural Philosophy and Mathematicks a great Branch of their Literature, and we see how universal Academick Knowledge has been approved of by the great Encouragement and large Priviledges the Schools have received from our Wilest Princes, and by the general agreement of the most sober and considerate men, in sending their Youth to receive their Education therein. This shews how universal Philosophick Learning has been approved; and the only Question that can be raised at present is, Whether the Natural Philosophy formerly professed in the Schools, or that which is at present prosecuted by the Societies lately Instituted in several the most noted parts of Europe, be the True Philosophy, or method of Investigating Nature. But surely this will be no longer a doubt, when we consider how unsatisfactory were the ancient Notions of Philosophy, which then consisted rather in Disputes, and Verbose empty Stuff, than in any Curious Discovery of Natures Actions. If a man could prove Pro and Con, whatever was proposed, and maintain

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tain this Dispute for two or three hours by vain Distinctions and idle Evasions, he presently gain'd the Name of a Philosopher, tho all the while he had no manner of Notions in his Brain answerable to those senseless words he threw out. I would fain know, what Notion any man has of Light upon pronouncing the definition of it, Actus Perspicui quatenus Perspicuum, does he hereby in the least understand any of the properties of Light, or how 'tis affected by Refraction or Reflection? or does this Definition lead him to improve Light for the advancement and help of our senses, or other advantage of mankind? And now, my Lord, if this kind of Philosophy which formerly fill'd our Schools were look'd upon as deserving the Favour of the Great, how much more shall the present useful Enquiries of the Ingenious deserve Incouragement? But that this may be made a little more evident (though I am perswaded 'tis sufficiently manifest already to Your Excellencys Discerning Judgment) I shall crave leave to be more particular in this matter; And I shall begin with one of the most Considerable and Universal Concerns of Mans Life, I mean Navigation, wherein the Philosophy of the Moderns has been very much Exercised; The Magnet by its Capricious Variations is at present in several parts of the World rendred almost useles; but diligent search is daily made after its Vagarys, that we may not be deprived of the unspeakable benefit of one of the most surprising Phenomenas in Nature. Certainly men might
have

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have Disputed from Age to Age of Sympathy, Antipathy, and the Occult Qualities of the Loadstone, before they should ever arrive to the Knowledge or Rectification of one of these Exorbitances, which diligent Experimental Philosophy presently discovers, and further Experiments may hope to Rectify. In relation likewise to Navigation are all the Enquiries into the Theory of the Tides, and the Moons motion, all the Endeavours for the Longitude, and the Diligence used to Rectify Astronomy. The Philosophy of Hydraulicks and Hydrostaticks tends to the great Emolument and Pleasure of Mankind, no place being habitable without Water, and Curious Fountains adorn the Gardens and Houses of the Greatest Princes. Agriculture, and the Philosophy relating to Earth and Vegetation, is of universal extent over the face of the World; and how mightily it has been lately prosecuted, is evident from the Incomparable Works of Mr. Evelyn and others. By the Doctrine of Light, and the Properties thereof, our most Noble Sense has been Improved to an Acuteness some thousands of degrees beyond its natural Abilities; and the Telescope and Microscope discover to us new Worlds and Animals, extending the Creation farther than 'twas possible for the Ancients to imagine. There is no state of Life that is not concern'd in Meteorology, or the Philosophy of the Weather. And all that have heard of the late contrived Baroscopes, Hygrosopes, Thermoters, &c. may plainly perceive what Endeavours have been made towards it, and their success.

I might

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I might proceed on, and fill a Volume with Instances to Your Excellency of the usefulness of an Active Experimental Philosophy, but I stop my hand, and shall only declare in short, that from this alone we are to expect advancements in the most important concerns of humane life; the Virtues of Animals, Plants and Minerals for the health of mans Body, Architecture Civil and Military for the pleasure and security of his life; all kind of Machines and Movements for the increase and due application of his strength, proceed from this kind of Philosophy, which consists in Actions, not in Words, and prosecutes and sometimes improves Nature by the same steps that she her self acts by. And indeed, my Lord, if we consider Nature as an Active Principle, we must needs acknowledge that she is to be prosecuted by Action, and not by Verbose Disputes; There is nothing that the mind of man will not find something to say for or against all day long; and therefore of the School-Disputes there is no end; but Experiment is matter of Fact, and strikes the Senses so forcibly, that there is no opposing it. Who is it that will now question the force of Gun-powder, or whether the Mercury rise and fall in the Baroscope at fair or foul Weather?

'Tis therefore this Philosophy (may it please Your Excellency) that deserves Favour and Encouragement from the Powerful and Wise; and as God has made all things in number, measure, and weight, that Learning which teaches us the affections

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 of number and measure, as being the Foundation on
 which true Philosophy is to be superstructed, is chiefly
 to be respected. There is no part of Philosophy where-
 in the Mathematicks are not deeply ingredient, and on
 them depend the Arts of War as the Delights of
 Peace, and even Sacred Theology it self is in a great
 measure beholding hereto for its help, if we consider how
 far the Chronology of the Holy Scriptures is help'd
 out by Astronomy, and the Doctrine of Eclipses.

And as I have presumed to detain Your Excellency
 thus long, in declaring how true Philosophy deserves
 Your Protection, so give me leave to add one word more
 in declaring how Your Excellency deserves the Tu-
 telage of true Philosophy. Your Illustrious Father,
 of Happy Memory, was no unactive Speculative Philo-
 sopher, but was zealous and forward in promoting the
 Noble Design of the Royal Society, as appears to his
 immortal Praise in their Incomparable History, and
 by their Electing him, and his willing Compliance to be
 their Protector. Neither did he only countenance Phi-
 losophy by his Authority and Favour, but actually
 drew the Sword in its defence against the great Sea Mon-
 ster that would have destroy'd the Virgin; as is mani-
 fest from his Learned and Ingenious Treatise against
 the Leviathan. All the World that know Your Lord-
 ship, must acknowledge that this and other his extraor-
 dinary Virtues are entail'd on Your Excellency, who
 has not buried the Talent in a Napkin, but improved
 it

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it to an high degree. Your Excellency has Honoured the Royal Society by being one of their Number, and their last years Council was compleated by Your being one of them. So that upon all occasions Your Excellency has shewn Your Self a zealous Patron of Learning and Ingenuity; and though I dare not on either score challenge Your Excellency's Protection of what I here present You, for I am afraid to trust my own Judgment in my own Concern; yet I will presume to implore Your Favour towards it as 'tis new, and not disliked by some Ingenious Men to whom I have shewn it. Whether it may deserve Your Excellency's Countenance for its Use and Advantage, I leave to Your own discerning Thought, after Your Lordship is pleased to consider what I offer in the First and Second Chapters. ~~However, this I will venture to assert before-hand,~~ that as Clocks or Time-keepers by Wheels and Weights have been an ancient Invention, (if we believe Severinus Boethius to be the first Author of them, 'tis above 1100 years ago) yet in these later days they have received a most high advancement (to the immortal glory of M^{rs}. Hugen) by the addition of a plain and simple Swag or Pendulum; since the days of H. zekiah, wherein we ~~first find~~ ~~as~~ mentioned, I have not heard of a more plain and easy addition for the advancement of Dials, and enlarging and ascertaining their Use, than what I here propose. And though I shall not be so vain, as to expect its sudden propagation, or any great Glory from the Contrivance;

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vance; yet if I can be so happy as thereby to Divert and
Please Your Excellency, and in some measure to pur-
chase Your Favour, 'tis the highest Ambition of

Your Excellencies

Dublin, April
17. 1686.

Most Obedient Humble Servant,

WILL. MOLYNEUX.

Sciothe-

Sciothericum Telescopicum,

OR,

A New Contrivance of Adapting a Telescope
To a large Horizontal Dial, for observing the
moment of Time by Day or Night.

CHAP. I.

The Use and Advantages of this Contrivance.

EVery one the least versed in *Astronomy*, does know of what great Concernment the Observation of the exact Moment of Time is therein. Without this no Celestial Observation can be performed Accurately, and *Astronomy* is like to receive but little advancements. Hereby the *Tables Astronomical* are approved or rejected, and Calculations found true or false. So that whatever advantages we propose to our selves by a Correct *Astronomy*, we shall find our present Indevour to be helpful towards them. And certainly there are some signal uses expected from it, since so many *Kings, Princes, States, and Learned Men* in all Ages, particularly our late and our present *Sovereign of Great-Brittain, The present French King, and formerly some Kings*

B

of

of *Denmark*, have been at such great expences both of *Time*, *Labour*, and *Charges*, for the advancement thereof. But I shall pass by all other excellent uses that are expected from an *Accurate Theory* of the motions of the *Heavenly Bodies*, and shall only insist on one particular, wherein the *Observation* of the exact *Moment of Time* does more immediately tend to the use of mankind. We all know how universal the whole *World* is concern'd in *Navigation* and *Commerce*, by *Ships* flying before the winds, and floating on the *Seas*; *Nation* converses with *Nation*, as every man with another by words flying in the *Air*; by this, *Civility*, *Learning*, and all *Politeness* is propagated; we are made acquainted with one anothers *Laws*, *Constitutions* and *Manners*, we mutually reap the *Fruits* of each others *Countrys*, and are no longer *Strangers*, but *Fellow-Citizens* of the *World*. And for all these advantages (at least for our more securely reaping these benefits) we are in a great measure, if not wholly, beholding to *Astronomy*. For whoever has but inquired into the first *Rudiments* of that *Science*, does very well know, how far the determination of the *Longitude* of places, and consequently the advancement of *Geography* and *Navigation* depends thereon. But whether this be attempted by *Eclipses* of the *Sun*, *Moon*, or *Stars*, or by the *Immersion* and *Emersions* of *Jupiter's* *Satellites* into or from his shadow, or by the *Pendulum Clock*, at *Sea*, &c. In
all

all these methods the *Observation* of the exact *Moment* of time is necessary, for otherwise the *Horary* distance of the place of *Observation*, and consequently the distance on the *Æquator*, or *Longitude*, from an assigned place shall be uncertain.

CHAP. II.

Of the Methods used for observing the Exact Moment of Time, and their Inconveniencies and Troubles.

THE common ways used by *Astronomers* for observing the *Time* are, either by *Dials*, or by taking the *Suns Altitude* by day, or by the *Altitude* of fixt *Stars* by night, taken by large and accurate *Instruments*; or by observing the *Altitude* and *Azimuth* of the *Sun* or *Stars*, or lastly by the *transits* of *Sun* or *Stars* cross the *Meridian*, or the coming of some of the *Circum-Polar Stars* in the same *Vertical* with the *Pole-Star*. All these methods have their inconveniences, or at least, are attended with far greater trouble than that which I shall propose.

And first for *Dials*, unless they be very large they will not admit of divisions so minute, as are requisite to the nice determination of the *Time*, as into minutes, half-minutes, and quarter-minutes. And when they are so large as to admit such, the uncertain shadow that is cast from a long

Gnomon renders them useleſs for ſuch niceneſs; for though your *Dial* have on it every quarter-minute, the moſt accurate eye cannot tell where the ſhadow determines to a minute. Another grand defect of *Dials* is, that they ſerve only in the day-time, and that too, when the Sun ſhines out intenſely. So that for the night (which is the chief time of *Aſtronomical Obſervation*) *Dials* are perfectly uſeleſs.

Perhaps I need ſay no more to recommend my preſent contrivance, then to aſſert that it clearly takes off both theſe inconveniences. For thereby the larger the *Dial* is, and the more minute the diviſions are, the more accurately is the Time obſerved. A long *Gnomon* in our way hinders not at all, nor is there any uncertainty from a *Penumbra*. And moreover 'tis adapted as well to ſerve by the Stars at night, as by the Sun in the day, and gives the hour almoſt by inſpection at a Star, as at the Sun. And withal has this advantage, that when the Sun is over-clouded, ſo as but juſt to be ſeen faintly, and caſts no ſhadow at all on a *Dial*; yet in our way, if he be in the leaſt perceivable by the eye, the Time may be exactly told by him.

I cannot ſay that there are inconveniences in obſerving the Time, by taking the Sun or Stars *Altitudes* and *Azimuths* by day and night; and yet if we reckon trouble an inconvenience, theſe methods are not free from it. For, beſides the difficulty

ficulty and charge of obtaining Instruments large and accurate enough for doing these, and the dexterity and long practice that is requisite for rightly managing them: This method is attended by the trouble of Calculating the most difficult of oblique spherical Triangles, viz. that wherein three sides are given to find an *Angle*; and this, if the observation be by the *Altitude* of the Sun in day-time; for by the *Altitudes* of Stars by night, the Calculation is yet more tedious and troublesome. And if so, what a toil must it be to suppute twenty or thirty of these in a night? Besides, the *Altitudes* of the Sun towards the *meridian* alter so slowly, that for a good while both before and after noon, 'tis not safe to rely upon them. This inconvenience indeed is something remedied at the Stars, where I can choose those that are of a proper *Altitude*, and conveniently remote from the *Meridian*, and the observation of the *Azimuth* remedies the other uncertainty proceeding from their vicinity to the *Meridian*. But, as I said before, the trouble of Calculation attends the observation either of *Altitudes* or *Azimuths*; whereas the Instrument I propose, does the matter with the greatest certainty, and greatest ease imaginable. There is requisite therein, besides a plain and simple observation, no Calculation by *Triangles*, or any other Operation, save only the addition and subtraction of two or three small numbers, to be had in the
Tables

Tables ready Calculated; and that too, only when the observation is by the Stars at night.

The methods of observing the Time by the *ap-pulses* of the Sun or fixt Stars to the *Meridian*, or the *Circum-Polars* coming in the same *Vertical* with the *Pole*, are plain, simple, and easy. We have the first described by *Riccioli* in his *Almagestum Novum*, *Lib. 5. Cap. 15. Prob. 8.* where he shews us how to perform it by his *Triangulum Filare*. The latter way is described by *Sir Jonas Moore* in his *Compendium Mathematicum*, p. 118. and more fully in the Royal Almanacks for the years 1676, 1677, 1678. And I have by me Tables laboriously and carefully Calculated by my esteemed Friend *Henry Osburn Esq;* an excellent Astronomer and Mathematician for *Dublin* and other *Latitudes* in *Ireland* to years lately past, and to come, which are to be used in this way. But, though both these ways (as I said before) are very plain and easy, yet they serve rather to rectifie Watches and other Time-keepers, then to shew the Time themselves throughout the whole course of an observation; as, suppose it were an Eclipse of the Moon, perhaps when a spot either immerses or emerges into or from the shadow, there is not at that very instant, and perhaps will not be for many minutes after, a Star in the *Meridian*, or under the *Pole-star*, to tell me the moment of that time.

But what I propose does as constantly (and not by

by fits) shew, and follow the time, if duly managed, as the Hand of a well-going *Pendulum-Watch* indicates the hour; that is to say, it tells you the present minute, and quarter of a minute, whenever you are pleased to look, as well as any other, past, or to come.

C H A P. III.

A Description of the Instrument.

THe contrivance of this Instrument consists in making a very large *Horizontal Dial*, adapted to your proper *Latitude*, and capable of receiving divisions into minutes, and parts of a minute, fitted with a large, strong, and double *Gnomon*; I call that a double *Gnomon* that casts the morning-shadow from its Western-edge, and the afternoon-shadow from its Eastern-edge, and the noon-shadow by its thickness. This *Dial* is to be furnished with two pair of Sights or Rulers; one to serve in the morning, or for Stars on the Eastern-side of the *Meridian*, t'other to serve in the afternoon, or for Stars on the Western-side of the *Meridian*. Each of these pair consists of two movable Rulers, one I will call the *Horizontal-Ruler*, t'other the *Gnomonick-Ruler*, or *Stile-Ruler*. These two Rulers are to be so adapted, that their two edges that are next the

the *Gnomon*, may be perpetually in the same plane with each other, and at the time of observation, that they both may be in the same plane with their correspondent edge of the *Gnomon*. On the *Stile-Ruler* are fixt Telescopick Sights, and the cross hairs in their due place. But all this will be made plainer by the first Figure, in which I shall represent that Instrument which by my directions was made in *London* in the Summer 1685 by *Richard Whitehead* Math. Inst. Maker, who took the directions from my self, but made the Instrument in my absence.

The first Figure represents the Instrument in Perspective, having one pair of its Rulers in a posture of observation; t'other pair lying carelessly on the plane of the *Dial*. 'Tis a large Octogonal Brass-plate, clear'd in the middle, except only the cross-bars *z z z* of the same piece with the Plate, left for strengthning the Instrument, and receiving the Stile, the Plate is of a moderate thickness about $\frac{15}{100}$ of an inch, the Diameter of the largest circle it receives is 18 inches, 'tis supported from the plane on which it stands about three quarters of an inch by three brass-feet *2. 2. 2.* the bredth of the limb for receiving the Figures and Divisions, is $2\frac{1}{2}$ inches. The thickness or bredth of the face of the Stile *malb* is $\frac{7}{10}$ of an inch. The Divisions actually expressed upon it are hours, half hours, quarters, five minutes, minutes, and half minutes, and

and one may easily judge of the third part of an half minute, that is, 10 seconds. Down along the face of the Stile (some perhaps will call it the back of the Cock) there runs a deep Groove, to receive the screw *m. l.* which screw, by turning the handle *n*, raises and lets down the Nut *op*, on which the Stile-Ruler *ef* rests, and is thereby raised and let down, as the Sun or Star requires. This Nut *op* is furnish'd with a Return'd Fork, meerly to keep the Stile-Ruler from an accidental fall, if any thing should chance to move it rudely. But I shall describe this Nut *o.p.* presently by it self. The screw *m. l.* is fixt both at top *m* and bottom *l*, so that it only is to be turn'd round by the handle *n*. the Stile is fixt most strongly by the screws, *q. r.* I come next to describe the Rulers, and first for the *Horizontal Ruler c d*, this turns most truly upon the Center of the hour-lines and divisions at *k*, and has so much of the root of the Stile clear'd off, as its thickness requires to permit its motion freely; the line *c d* of this Ruler crosses the Center most exactly, and the edge *c d* is neatly cypher'd off (as the Workmen call it.) Wherefore the right line *a b* of the Stile, and the line of the under-edge of the *Horizontal Ruler c d*, meeting, and crossing perpetually in the same point *k*, they shall be always in the same plain, (by *Prop. 2. Eucl. 11. Book.*)

To the other end of this *Horizontal Ruler* there

is adapted the Stile-Ruler *ef*, which by means of the joynt *xy* (which I shall describe by it self) obeys two severall motions, *viz.* one upwards and downwards, as govern'd by the Nut and screw on the Stile; and t'other Eastward and Westward, according to the *Azimuth*, or as it follows the *Horizontal Ruler*. And in these two motions, 'tis so contrived, that the line *cd* and *ef* may perpetually cross each other in one and the same Center at *y*. Hence it comes to pass, that the right line *cd*, and the right line *ef*, are always in the same plain. Wherefore laying the line *ef* to the edge of the Stile, the three lines *ab*, *cd*, *ef*, are in the same plain; and consequently directing the line *ef* to the Center of the Sun, the edge of the *Gnomon*, and the edges *cd* and *ef* of the Rulers are in the plain of the shadow; and therefore the Ruler *cd* cuts the division on the limb of the Dial-Plate, where the shadow would cut it.

To the Stile-Ruler are fixt, the Eye-glass at *b*, the Ring that carries the cross hairs at *g*, and the Object-glass at *i*; The manner of adjusting all which, I shall shew presently.

But I must not forget to mention the Artificer, whose hands were imploy'd in this Instrument, Mr. *Richard Whitehead* living in *Gunpowder-Alley* in *Shoo-Lane*, by *Fleetstreet*, *London*; a most exact and careful Workman, of whose skill and curiosity in making

making Mathematical Instruments, I have more than this one instance; And I dare venture to recommend any Gentleman to him, as a most exact performer.

CHAP. IV.

Of the Stile-Nut, and Ruler-Joynt.

IN the foregoing Chapter I have promised a particular description of some of the fore-mentioned parts belonging to this contrivance; and first of the Nut *op* (*Fig. 1.*) This I have represented by it self in *Fig. 2.* *ss* is the Stile-screw passing through the Nut *n*, which is lined with a female-screw; this Nut slides in the Groove along the face of the *Gnomon*, so that the whole thickness of the screw lies under the said face. *pp* are two arms that clasp the Stile to make the motion of this Nut more even and steady. *f a b* is the Return'd Fork to keep the Rulers, when they are supported on this Nut, from any accidental fall, and is very useful when the *Horizontal* Ruler is removed far from the *Meridian*. And least at any time, either of the Returns *a b*, may stand in the way of either of the Rulers motions, the parts *a b* are movable at *a* by a riveted joynt, and may be turn'd clearly downward out

of the way, as at *ac, ac*. At *a* there may be fixt a small gentle spring, that may lightly bear the edge of the Ruler constantly to the edge of the Stile; but this contrivance is not on my Instrument, though it may be easily added.

The several parts of the Ruler-joynt (which indeed is the most curious piece of Mechanicks about the whole Instrument, and on which the whole affair does chiefly depend) are represented all apart in Figures by themselves. *Fig. 3.* shews the outward end of the *Horizontal Ruler*, *aaa* are the holes that admit the screws which fasten it to the other parts of the joynt, particularly to the socket shewn in *Fig. 5, 6, 7, 8.* at *aaa*. *Fig. 4.* shews the other end of the *Horizontal Ruler*, where it turns on the Pin in the Center of the Dial at the root of the Stile (at *k Fig. 1.*)

Fig. 5. expresses the appearance of the joynt, when discover'd of *Fig. 3.*, *ibi* is a perfect hinge, the Center of whole motion is in the line *fb, iii* that part of the hinge belonging to the Ruler, *bhb* that part belonging to the Tumbler or Nut expressed in *Fig. 7. bbb. aaa* the socket wherein the Tumbler *bhb* moves round and truly on the Center at middle *b*, which also is in the Center of the hinges motion. This socket *aaa* is represented in *Fig. 8. bb* are the screw-holes for the Ring and Snout of the Eye-glass.

Fig 6 shews the under-side of *Fig. 5.* where the
same

same parts are marked with the same letters; *c* is the round Pin of the same piece with the Nut *b b b*: This is more plain in *Fig. 7*, this Pin turns truly in the hole of the socket *c* in *Fig. 8*. whilst the Nut *b b b* of *Fig. 7*. rests and turns on the bottom of the said socket *b b b* *Fig. 8*. This Nut is kept close in the socket, and there made to move steadily and truly by the Pin *e* and springing Plate *d* *Fig. 6.* and *9.* which Pin *e* passes over the Plate *d* through the Pin of the Nut *c*.

CH A P. V.

Of the Telescopick Sights, and their true Adjusting.

FOR the true adjusting of the *Telescopick Sights* *b g i* (*Fig. 1.*) to the Stile-Ruler *ef*, there are three things requisite. But first I must advertise, that the Stile-Ruler *ef* ought to be so strengthened, that as it rests on the Nut *o p*, it may not at all bend, but that the line *ef* may be one right line as nigh as possible. But in this particular the greatest niceness is not requisite, for if the Ruler chance a little to deviate from its exact straightness, I shew a way of rectifying the error by the cross-hairs or *Mensurator*.

The first thing requisite to the placing of the *Telescopick Sights* is, that the ring *g*, that carries the

the cross-hairs or other *Mensurator*, be at a due and exact distance from the Object-glass *i*. This distance is to be exactly the total length of the Object-glass; but that we may get this total length more accurately than by any admeasure-
ment, the Eye-glass and Object-glass being adapted together, and placing the *Mensurator* in the ring *g* between the Object-glass and Eye-glass at the distance of the *Focus* of the said Eye-glass, let us look at some object distant 4 or 5 miles, and moving your eye before the Eye-glass, observe whether the *Mensurator* seems to move upon the said object; for if it do, then the *Mensurator* is not at its exact distance from the Object-glass, but the Object-glass is to be removed farther or nigher, till the eye looking at such a distant object, and moving before the Eye-glass, perceives the *Mensurator* as it were fixt and immovable on the object. If in raising the eye, the object seem to fall down on the *Mensurator*, or in depressing the eye, the object seem to rise on the *Mensurator*, then is the Object-glass too nigh the *Mensurator*. But if in raising the eye, the object seem to rise on the *Mensurator*; or in depressing the eye, the object seem to fall on the *Mensurator*, then is the Object-glass too far from the *Mensurator*. All which will be evident from the 10 Fig. wherein let *AB* be a distant object, whose middle point *C* is projected by the Object-glass *D* at *k*, let *m n i* be the *Mensurator*

furator to which the Object-glass is too nigh, and *m n 2* the *Mensurator* from which the Object-glass is too far, *e, f, g*, the eye placed at three different postures; in case of the first *mensurator*, if the eye rise from *e* to *f* it perceives the point *k* depressed from *1* to *b*; or if the eye fall from *e* to *g*, it perceives the point *k* raised from *1* to *l*; and here the *mensurator* is too nigh the Object-glass. But in case of the second *Mensurator*, if the eye rise from *e* to *f*, the point *k* seems to rise on the *Mensurator* from *2* to *r*; or if the eye fall from *e* to *g*, it perceives the point *k* fallen from *2* to *s*, and in this case the *Mensurator* is too distant from the Object-glass; but if the *mensurator* be exactly in the *Focus* at *k*, let the eye rise or fall, the *mensurator* seems fixt and steddily upon the object. And this is the first thing requisite to the adjusting of these Sights.

This Affair is usually so well adjusted by the Workmen, and when once adjusted, is never alterable, that I need shew no way of providing for it. But in short, it may easily be provided for, by making the holes for the screws, that fixt the Object-glass-ring to the Stile-Ruler, long slits, so as to slip forwards and backwards on the necks of the screws; and when the Object-glass is at its right posture, the screw-heads may pinch and fix the Ring steddily.

The next requisite is, that the line of *Collimation* from the *mensurator* through the Object-glass, and
so,

so to the object, run parallel to the line *ef* of the Stile-Ruler. This *Parallelism* consists in two manner of ways, first, that the *mensurator* be neither more to the right or left-hand than requisite; or secondly, that it be neither higher or lower than it ought. The first error makes the *Azimuths* be shewn false, and the second error makes the *Altitudes* be shewn false: but an *Horizontal Dial* tells the time by both *Altitude* and *Azimuth*; therefore both these are to be taken care of.

And first for the Fabrick of the Ring *g*, that carries the *mensurator*; this is so to be order'd, that first the Ring may be moved more to the right or left-hand, and there to be fixt; and secondly, that the *mensurator* may be raised or depressed according as is requisite, and there fixt. The first is obtain'd by making the holes for the screws, whereby this Ring *g* is fixt to the Stile-Ruler, long slits, so that they may slip on the necks of the screws before the Eye-glass, and when they are in their right posture, may be pinched close by the screw-heads, and there fixt. Thus therefore we give the *mensurator* a motion to the right or left-hand, by which 'tis brought to shew the *Azimuth* right, as I shall declare presently. Next, that the *mensurator* may be raised or depressed, instead of cross-hairs in the Ring *g*, let there be placed therein a very strong Steel-needle, that may end in a most fine slender point: this needle

is to be a screw almost its whole length nigh to its smaller end, and to be screw'd through the top of the Ring *g*, so that its smaller end may pass through the center of the Ring; by screwing or unscrewing of this Needle, we depress or raise its fine point in the Ring. And thus much for the contrivance or Fabrick of the Ring.

I come now to shew the way of managing these two motions in the *Mensurator*, so as to bring it to its true posture, and there fix it. And first for rectifying the *Mensurator* so as to shew the true *Azimuth*. There are two manner of ways for doing this. The first is more troublesome and tedious, but being most universal to the fixing of *Telescopick Sights* on all *Rulers*, I shall here describe it. In *Fig. 11.* let *ABCD* represent a *Ruler*, whose edge or side *AC* is exactly parallel to its side *BD*. To make a *Ruler* thus parallel, is not difficult to a good Artificer; but I shall shew how to try whether these sides are parallel or not. *E* is the Object-glass, *F* the *Mensurator*, *G* the Eye-glass. On a plain Board strike two round Brass-pins, as suppose at *H* and *I*, to these apply the side of the *Ruler BD*, so as it may rest against the pins, which are therefore not clearly buried into the Board, but stand out about $\frac{1}{10}$ of an inch: then look through the Glasses, and observe where the *mensurator* falls on an object distant a mile or two. Then remove the *Ruler*, and apply its other

D
side

side *AC* to the other side of the pins *HI*, and observe whether the *mensurator* falls on the same point of the object as before. If it do so, then are the sides of the *Ruler AC, BD* parallel; if not, then the sides are not parallel. The reason that so remote an object must be chosen, is, that the breadth of the *Ruler* may subtend an inconsiderable *Angle* in a circle, whose *Radius* is the distance of the object. Having found the sides of the *Ruler* to be parallel, the next thing is, to make the line of *Sight*, or line of *Collimation LK* parallel to these sides, for therein consists the first rectification of our *Stile-Ruler*. The method of doing this is much the same with what immediately precedes, but requires another disposition of our pins; for now we are to raise our plain Board so as to stand edge-wise, nigh perpendicular to the *Horizon*, and the pins must stand an inch or more out from the Board, almost parallel to the *Horizon*; then resting the under-side of the *Ruler* on the pins, and applying its edge to the plain Board, observe the point of a remote object whereon the *mensurator* falls; in this posture the Glasses stand uppermost, or on the upper-side of the *Ruler*, and over-look the pins. Then remove the *Ruler*, and hang the other side thereof on the pins, that now the Glasses may be on the under-side the *Ruler*, or under-look the pins, and resting the edge against the plain Board, observe whether

ther the *mensurator* fall on the same point as before; if it do, then is the line of *Collimation* *LK* parallel to the sides *AC*, *BD*; if it do not, but falls to the right-hand apparently of the said object, then is the *mensurator* to be removed (by means of the screws and slits in its Ring) to the left-hand: the contrary requiring the contrary. And thus by frequent repetitions and trials we at last bring all to rights. The most convenient Board or Table for this Operation, is a Surveyors Plain-table; for these being usually made true, and readily and steddily obeying all motions and postures, and are easily fixt thereat; they are to be chosen before any other. Note also that the pins to be used with this or such like Table, are to be strong brass-wire, which having its roundness from its drawing, is sure to have its sides parallel. By this method, the line of sight of any *Cylindrical* or *square Telescope* may be made to run parallel to its sides, for finding the *Declination* of the *Magnet* according to the methods lately proposed by *Monsieur Hautefeville*, and *M^{n^r}. Sturmius*, in the *Journal des Scavans* 23. Aug. 1683. and in the *Acta Lipsiæ*, An. 1684. mens. Decemb. and for want of this method, what *M^{n^r}. Sturmius* says in the foresaid *Act. Lips.* pag. 579. is very defective. For thus he, *Sola tubi locatio ut axis visionis per medias lentes excurrens meridiana lineæ exacte respondeat difficultatis quippiam habere videbatur, verum & huic infirmitati præsens, uti*

credo, inventum est remedium, &c. And the remedy he tell us is, that the *Tube* be made a parallelepiped of Wood or Brass; for then, says he, applying the side of your *Tube* to the *meridian line*, the *Axis of vision* will be parallel to the said *meridian line*. But I must crave leave to deny this, unless first it be rectified, so that this *Axis* runs parallel to the side of the *Tube*, which he was not aware of: and this is to be done by the method I have just now proposed, which serves likewise in many other Occasions, Experiments and Practices, wherein a *Ruler with Telescopick Sights* is requisite.

The second method for effecting this said rectification, is more easily applicable to our *Dial*, and withal is sufficiently accurate, as doing the business to 10" or 15" seconds of time at utmost, and by a careful and curious eye, may do it to half that or less. Place the *Dial* so before the Sun, that as high as possibly the eye can judge, the *Gnomon* may be equally enlightned on both sides, that is, that the shadow of the *Stile* may fall exactly on 12 a Clock, or the *meridian line* of the *Dial*. This perhaps will be said by some to be no fair proceeding, because in this we cannot tell where the shadow does exactly determine, and that this being one of the inconveniences our *Dial* pretends to remedy, if we rectify our *Dial* by supposing we can determine that; & if afterwards we determine that

that by supposing our *Dial* rectified, it will seem a circle in *Argumentation*, but yet I say, they that will try it, will find it otherwise; for I can place the *Dial* so, that the shadow from the thickness of the *Stile*, or at least the *Penumbra* from both the *Stiles* edges, shall so equally fall upon the 12 a Clock-lines, that 10th seconds of time, or less, shall sensibly alter the equality. And 'tis not the same case in telling when this shadow comes to an equality on both the 12 a Clock-lines, as in telling when the shadow of one of the single edges of the *Gnomon* comes to any of the other hour-lines. For in the first case I am only to judge by comparison, when the shadow is come as much to one of the 12 a Clock-lines as to th'other; but in the latter case I am to judge positively without comparison. Wherefore having the *Dial* in this posture, bring the *Horizontal* and *Stile-Rulers* just to 12 a Clock on the *Dial-Plate*, and observe whether your *Mensurator* divide the Sun equally into an *Eastern* and *Western* half; if it do so, then is your *Mensurator* right; if not, the *mensurator* is to be moved something to the right or left hand, as is requisite, till at last by frequent repetitions and tryals we obtain our desire. And thus much shall suffice concerning the rectification of the *mensurator* to the right or left hand, so as to shew the *Azimuths* truly. I shall only take notice, that when one pair of the *Rulers* is rectified, the other pair

pair is easily rectified by them; for bring the rectified pair to the 12 a Clock-line on the *Dial*, and move the whole body of the *Dial*, till you get the center of the Sun, or any Star on the *mensurator*, then immediately bring the unrectified pair of *Rulers* to the 12 a Clock-line on their side, and if the Sun or Star be not exactly on the *mensurator* on this pair, the *Mensurator* is false, and must be rectified, as is requisite: but this being plain to the meanest capacity versed in these matters, I shall insist no longer thereon.

But I proceed to the other Rectification of the *Mensurator* for rightly telling the *Altitudes*, as I have said before. And tho there be not so very great exactness required herein, especially in telling the time by the Sun or Stars when nigh the *Meridian*; yet a gross error herein must not be allow'd. This likewise is to be performed much after the same method with the former Rectifications, for 'tis but contriving some way for inverting the *Ruler*, so that its errors of this kind may be perceived, as by our former Inversions we discovered the errors of another kind. But to make this a little more plain, this error consists in the line of Sight not running parallel with the under-side of the *Ruler*; wherefore if we have (suppose) a *Deal-board* whose two faces are exactly parallel, and this *Board* be raised and fixt
so

so as to lye steddily along one of its edges, and on each face of the Board there were a ledge to support the Ruler; and the Ruler be applyed by its under-side to one face of the Board, and the object marked out by the *Mensurator*, diligently observed; and then the Ruler applyed by the same under-side to the other face of the Board: If the same object be now cut as formerly, ly, all is right; if not, the *Mensurator* (being such as I have formerly described) must be screw'd or unscrew'd, so that the fine point thereof may be lower or higher in the Ring, according as is requisite.

Another more easy way of discovering and rectifying this Error to sufficient accuracy, is thus: After you have by the former methods brought the *Mensurator*, and thereby the *Ruler*, to shew the true *Azimuth*; observe the time by the *Dial* when the Sun is very far removed from the *Meridian*, as early in the morning; and at the same time observe by a good *Pendulum* Clock the hour, minute, and second; and note the difference between the *Dial* and Clock. Proceed thus to make an observation every quarter of an hour through the whole morning, still noting the differences: And because we may well suppose that a good *Pendulum* Clock in so short a course of time goes equally, these differences should be always equal; but if they are unequal, then the *Mensurator* shews the

the *Altitudes* fallibly. As for instance, suppose I find that approaching towards Noon these differences decrease, then does the *Mensurator* on the *Telescopick Ruler* shew the *Altitudes* too little, and therefore must be order'd so as to shew them more: if I find them increase, the contrary is to be done. The reason of this is plain, for an *Horizontal Dial* shews the time as well by the *Altitudes* as *Azimuths*; but as the time approaches towards Noon, the *Altitudes* are less concern'd in the matter than the *Azimuths*; and so less and less, till just at Noon the time is not at all shewn by the *Altitude*, but solely by the *Azimuth*. And consequently, if there be any error in the time arising from a false *Altitude*, it will appear by comparing two times or more together; some, wherein the *Altitudes* are much concern'd, and others, wherein the *Altitudes* are little or not at all concern'd; but such times are, early in the morning, and towards noon.

CHAP. VI.

The Way of Setting this Dial.

TO the true Setting of this *Dial*, there are two things requisite; First, that the plain of the *Dial* be in an exact *Horizontal* posture, this
is

is easily obtain'd by accurate levels, which are so common, that I shall mention nothing further hereof. The second requisite is, that the *meridian* or 12 a Clock-line of the *Dial* be in an exact *meridian* line. This indeed is one of the chief particulars that we are to take care of, for thereon depends the accuracy of the whole. But that we may not over-turn or neglect the whole affair for the difficulty of this one particular, I shall shew that this is no such insuperable hardship as may be imagin'd. Let us therefore consider, that what is here alledged as a difficulty in this *Dial*, may as well be made against the large *Azimuthal Quadrants* used by *Astronomers*, and their observations of *meridional Altitudes*, and *transits* through the *Meridian*; *Ricciolis Triangulum Filare* described *Almag. Nov. Lib. 5. Cap. 15. Probl. 8.* and all Enquiries after the *Declination* of the *Magnet* are to no purpose; for these wholly proceed on the discovery of an exact *meridian line*. And indeed there are ways sufficiently accurate described by several for finding the true *meridian*. Amongst others, *Hevelius* in the first Part of his *Machina Caelestis*, Chap. 16. has many contrivances: And I shall presently set down an Instrument and way of my own for doing it, inferior, I think, to none. But let us a little consider, how much a *meridian* line must be erroneous, to make a true *Dial* apply'd to it to err a minute in time; and this in our *latitude* of *Dublin*

will be 12¹ minutes of a *Quadrant* to make the *Dial* err a minute in time about noon; and for so much error in the *meridian* line the *Dial* shall err less than a minute in time about 5 and 6 a Clock, as is manifest to those that understand calculating hour distances for an *Horizontal* plain. Now 'tis a very rude way indeed that will not take a *meridian* line more accurately than to half or quarter 12 minutes error.

But last of all, (clearly to take off this difficulty,) if it be allowed that there is any most accurate way of telling the time of day or night, as by the *Altitude* of the Sun or Stars taken by large and curious Instruments, and Calculating thereby; I say, by such a way as this I will set my *Dial*, and then surely it will be granted to be in a true *Meridian*. And perhaps this very hint may shew as accurate a way as any in the World for finding a *meridian* line; for the 12 a Clock line of a large and true *Dial* that is Set by such an accurate observation must needs lye in a true *meridian* line.

But I am mindful of my promise, *Viz.*

CHAP. VII.

For Finding a Meridian Line.

THis is perform'd by means of the Instrument represented *Fig. 12.* *ABC* is a *Triangle*. (That which I have made is of Wood, but a good Mathematical-Instrument-maker would make one much better of Brass.) *AB* I call the *Perpendicular* side, *BC* the *Horizontal* side; in the end of the *Horizontal* side at *E* there is a screw, and from the said *Horizontal* side there strikes out a short arm *D*, in which also there is a screw; this arm *D* keeps the Instrument steady from tottering. By means of these two screws at *D* and *E*, the side *AB* is brought to its exact perpendicularity; but especially by means of the screw *D*, the edge *xz* of the side *AB* is brought to stand exactly in the same *Vertical* plain with the edge *mn* of the side *BC*, and this by help of a Plum-line. *F* is a plate of Brass, having in it a center-hole in the line *m, n* continued. *G* is a round, slender, but strong Brass-pin arising perpendicularly from the face *mn*. 1, 2, 3, 4, 5, &c. are round Brass-pins that rise and stick out from the face *xz*, at what distances we please. This Instrument being placed truly on an exquisite *Horizontal* plain,

so that its end *E* may something over-stretch the plain, and turning steddily on a pin in the center-hole *F*, lay a *Ruler* adapted with *Telescopick Sights* over the pin *G*, and any of the other 1, 2, 3, &c. as is found most convenient and agreeable to the *Altitude* of the Sun or Stars you observe by, at convenient times both before and after Noon. 'Tis therefore requisite that in the under-side of your *Telescopick Ruler* towards the end next the eye, there be fixt a pin, that resting on the pin *G* may keep the *Ruler* from slipping down in this declining posture. Let us then suppose that about 9 a Clock in the morning on the Summer Solstice, the *Telescopick Ruler* resting on the pin *G*, and the third pin in the perpendicular side takes the Suns *Altitude* exactly, then on the *Horizontal* plain draw a line along the face *m, n*. Again, let us suppose that the *Ruler* removed upwards to the fourth pin, and there resting, takes the Suns *Altitude* the same morning about half an hour after 9; then by the face *m n* let us draw another line on the *Horizontal* plain, and so let us proceed to elevate the *Ruler* on other pins, to make observations at 10, at half an hour after 10 the same morning, still drawing lines on the *Horizontal* plain at each observation.

Then again in the afternoon let us descend by the same steps or pins, by which we rise in the morning, diligently observing when the Sun comes

comes to the corresponding *Altitudes*, and drawing lines as before. Here we shall have three or four observations made in the morning, and as many in the afternoon. Then *bisecting* these *Angles* from the point *F*, if all their *bisections* are coincident in the same right line, we are sure that line is a true *meridian line*.

What is here said of the Sun, and of observing by it at the Summer *Solstice*, may be accommodated to the taking a *meridian line* at any time of the year by some *fixt Stars*, that being removed about 2 or 3 hours from the *Meridian*, have *Altitude convenient*; though in choosing of a great *Altitude* there is no great nicety; for if *Refraction* do interpose in the morning-observation, it interposes as much in the afternoon-observation; so that 'tis as if it did not interpose at all. And here we have a way of finding a *Meridian line* by the Stars at night, which is of no small advantage, they being not subject to sudden alteration of their *Declination*, and consequently this method may be practised at all times of the year.

And indeed if we are careful to have all things exact, *viz.* an exact *Horizontal* plain, the face *xz* in an exact *Vertical* with the face *mn*, the *Telescopick Ruler* and its Sights exactly adjusted by the methods in the foregoing 5th Chapt. and be very diligent and accurate in observing the true *Altitudes*, this method of finding a *meridian line* will
appear

appear inferior to none that has yet been proposed.

Such a plain *Telescopick Ruler* as is expressed in the 11th. Figure, of a convenient length, is sufficient in this practice.

CHAP. [VIII.]

The Manner of observing the Time, and exactly determining it by the Sun or Stars.

ALL things being rightly adjusted, and the *Dial* placed in an exact level or *Horizontal* posture, and by a true *meridian* line, look at the Sun through the *Telescopick* or *Stile-Ruler*, and bring the *Mensurator* upon the Suns center, then shall the *Horizontal Ruler* cut the hour, minute, and part of a minute most exactly. But for Setting, or finding the error of a Clock, the best way is to bring the *Horizontal Ruler* to some full Division, as, to some compleat minute; and by rightly managing the *Stile-screw* and *Nut*, and *Stile-Ruler*, observe when the center of the Sun and *Mensurator* come together, for that is the exact time to which you placed the *Horizontal Ruler*. And indeed through the *Telescopick Sights* we shall perceive the motion of the Sun so very quick, that we may determine its being on and off the *mensurator* to 2 beats of a second

second *Pendulum*. So that if it be granted me that I can bring and settle the *Horizontal-Ruler* to a full division exactly, and not to err in placing it thereat over 3, 5, or 7 seconds, I can determine the time of day or night to 3, 5, or 7 seconds. But if this great exactness will not be allow'd me, I say it is not on the account of any fault in the *Theory* of this contrivance, which I dare assert to be most accurate in it self, but on the account of the workmanship or deficiency of some of its parts. And that strikes not at the Inventer; let those that use them, and adjust them, look to their trunefs in all particulars.

The way of using this *Dial* at the Stars by night is much the same, only for these there are requisite the Tables, at the end of this Book, of the Sun and Stars temporary *Right Ascensions*. In looking at, or observing a Star through the *Telescopick Ruler*, the *Horizontal Ruler* cuts the said Stars *horary* distance from the *Meridian*, but then the hours are to be counted by the smaller Figures on the in-side of the limb. Thus, suppose I observe a Star, and find the *Horizontal Ruler* cut the ten a Clock line to which the great Figure of *X* is affixt, the *horary* distance of that Star from the *Meridian* is 22 hours, that is, twenty two hours are elapsed since that Star was last in our *meridian*, though really the Star be but 2 hours to the East of our *meridian*. This premised, the great Rule
for

for telling the time by the Stars is this; To the Stars horary distance from the meridian, add the Stars temporary Right Ascension, and from the Sum subtract the Suns Right Ascension, the remainder (rejecting 24 hours if need be) gives the hour, minute, &c. of the night.

I shall declare this by an Example. Anno 1686. April the 5th, between 9 and 10 a Clock at night I desire to know the exact time. I observe by *Spica Virginis*, and find (suppose) its horary distance from the meridian 22 h. 03' 15". The temporary Right Ascension of that Star for that year is, 13 h. 08'. 46", this added to the foresaid horary distance, the sum is 35 h. 12'. 01". At the same time the Suns place is γ . 26° 13' 59". Therefore its Right Ascension in time by the following Tables, is 1 h. 37'. 17". this subtracted from the foresaid sum (and rejecting 24 hours) leaves 9 h. 34' 44". the exact time of night. The reason and demonstration of this method depends on Astronomical Principles, which at present I must not undertake to illustrate, but are obvious enough to those that are versed in that Science, and will be very plain to those that consider the way of Calculating the hour of the night by the *Altitude* of a Star given, according to the most Learned and Ingenious *Pere Tacquets* Illustration thereof in his *Astronomy*, Lib. 5. Cap. 4. Num. 45.

And though there are other ways of ordering the foregoing *Data*, for getting the hour of the night,

night, yet what I propose is as plain and easy as any, and less embarras to the mind.

But here I must not omit one particular, and that is, that by our present contrivance, the *Right Ascension*, or place in the *Æquator* of any Star, is most easily and accurately obtained. How difficult and troublesome 'tis to obtain the *Right Ascension* of Stars, is known to Astronomers, and will appear to those that consult the forementioned *Tacquet Astronom. Lib. 5. Cap. 1. num. 4, 5.* but by this Instrument, and a truly rectify'd *Pendulum Clock*, the business is easily perform'd: For, from the sum of the Suns *Right Ascension*, and hour of the night (known by the Clock) subtract the Stars hour on the *Dial*, the remainder is the Stars *Right Ascension* in time, which converted into degrees and minutes, shews the *Æquatorian* distance of the Star from the first point of γ , or its *Right Ascension*.

CHAP. IX.

Of the Tables of the Suns and Stars Temporary Right Ascensions.

THE Table of the Stars *Right Ascensions* in time is plain enough of it self. 'Tis computed to the year 1686, and will serve for ten years to come without the error of half a minute.

note. 'Tis according to Riccioli's Catalogue of fixt Stars in his *Astronomia Reformata*. It consists of 5 Columns, the first contains the *Magnitudes*, the second shews the *Greek Letters* by which each Star is marked in *Bayer's Uranometria* Printed at *Ausburg*, 1603. The third has the numbers by which each Star is noted in *Ticho's Catalogue*, which is followed by most Authors, but is to be found particularly in *Tichonis Brahe Astronomice Instaurata Progymnasmatum*, pag. 258. and at the end of the *Rudolphine Tables* published by *Kepler*. The fourth Column contains the Stars names, as they are in *Bayer's Uranometria*; and also after each name we express the common name in *English*, and as they are to be found on our common Globes. And here we must not wonder to find that called *Right* or *Left* in the *Latin* names of *Bayer*, that is called *Left* or *Right* in the *English* common names; as for instance, *Herculis Humerus Dexter*, is in *English*, the *left* shoulder of *Hercules*. For the reason of this will appear plain to those that consider the Constellations of *Bayer*, and the common Constellations on our Globes. And I thought it not amiss to take all this care, and express all these marks for the sake of those that are not so well versed in the appearance of the Firmament, that they may the readier find out, and not so easily mistake any of the forementioned Stars. The fifth and last Column expresses

tes the *Right Ascension* in time of each *Star* to the present year 1686.

The Tables of the Suns *Right Ascension* in time require the Suns place to be had from some good Tables or *Ephemerides*. They consist of four Columns, the first contains every degree and 10 minutes of the Suns place, the second contains the hours appropriated to the signs between the *Vernal* and *Autumnal Equinox*, the third shews the hours appropriated to the signs between the *Autumnal* and *Vernal Equinox*; and the fourth and last shews the parts of an hour, that is, minutes and seconds, that are common to the opposite signs. By the side of each division are expressed the differences between every 10 minutes of the Suns place, for more ready making proportion. For Example, suppose the Suns place were γ $25^{\circ} 10'$ I find its *Right Ascension* 1 h. $33' 15''$. But if his place were α $25^{\circ} 10'$ the *Right Ascension* would be 13 h. $33' 15''$. Where we see $33' 15''$ common to both places. But suppose I find by a good *Ephemeris* the Suns place about my required time to be γ $25^{\circ} 15'$ then by making proportion I find by the Tables the *Right Ascension* to be 1 h. $33' 34''$. For I say if $10'$ (the common difference of the minutes) give $38''$, what shall $5'$ (the difference between $10'$ and $15'$) give, and it will be $19''$, which added to 1 h. $33' 15''$ makes 1 h. $33' 34''$.

I know the common way for giving the Suns

Right Ascension is by making a Table of the days of the month, and giving the Suns *Right Ascension* to the days. But this way, if it be *general*, and not appropriated to a certain year, is not accurate to a minute or two : and if it be *adapted* to a certain year, 'tis not *general*, and will not serve another time. Whereas our Tables of the Suns *Right Ascension* are *general*, most accurate to a second, and perpetual.

CHAP. X.

*Concerning the Astronomical Equation of Time,
and the Tables thereof.*

BEing now upon the business of Time, and the accurate observation thereof, so as thereby to regulate curious Time-keepers; it will not be improper to our subject to speak something of the *Inequality* of natural Days; a matter that has exercised the thoughts of all Astronomers in all ages: And though all have allowed that there really is such an *Inequality*, yet they have much disagreed in assigning its quantity, or demonstrating the reason and affections thereof; till at last our most Learned and Ingenious *English* Astronomer, and my Honoured Friend Mr. *John Flamsteed* *Math. Regius*, has determined the Controversy, and

and by most evident demonstrations has put the matter beyond further dispute, clearly evincing both the Reasons, Affections, and Quantity of this *Inequality*. His *Dissertation* concerning this is annex'd and publish'd at the end of the *Opera Posthuma Jeremiae Horroxcii*, Lond. 1673. 4^{to}. From which (with my esteemed Friend's leave) I shall present the Reader with the following Schemes and Demonstrations.

On account of the Sun's *Excentricity* from the center of the Earth's *Annual Orbit*, the *Diurnal* motion of the Earth is sometimes faster, and sometimes slower than the *mean* motion, and consequently the *apparent* Day is sometimes longer, and sometimes shorter than the *mean* Day. Which *Inequality*, and the Quantity of the difference of the *equal* or *mean* Day from the *apparent*, is thus demonstrated from the 13 Figure, according to the *Copernican System*.

Let $ABPN$ be the great *Orbit* in which the Earth is yearly carried about the *Sun*, the center hereof is C , A the *Aphelion*, or the Earth's place at Noon on that day that it is in its *Aphelion*, suppose the 18th. of *June*. B the Earth's place at Noon the day following. AL an assigned *Meridian* of the Earth. The arch AB , or the angle ACB , the mean motion of the Earth from the Noon of a given day to the Noon of the day following. L a point in the given *Meridian* turn'd to the *Sun*; which point,

point, whilst the Earth is carried in its Orbit from *A* to *B*, is rould by the *diurnal circumvolution* of the Earth from *L* through *O* in the first place *A* to *d* in the second place *B*; to which place, when the said point arrives, 'tis manifest that the Earth has performed a compleat revolution about its own *Axis*; because the *meridian B d*, in this its second posture at *B*, is made parallel to *AL* its yesterdays posture at *A*. But it is not yet apparent Noon, till the same point of the Earth by its revolution be brought to *e*, where 'tis turn'd directly opposite to the *Sun*, who governs the Civil Days. And that this time is not the same with the *Celestial* or *equal* Noon, will be proved, not only because the Earth has not yet performed its *mean* motion above its revolution, (tho this were a sufficient argument) but also because the *diurnal* motions about the *Sun*, and consequently the returns of any certain *meridian* to him, are very *unequal*; neither can they possibly be *equal* in respect of any point about which the Earth is not carried *equally*, as is sufficiently manifest from the inspection of the *Scheme* only. Wherefore the mean Noon and equal Time respects the point of the mean motion (*that is the center of the Orbit C*) and in our present instance is then when the *meridian* carried from *e* arrives at *f*, where 'tis directly opposite to the center of the *Orbit C*. And when it has gained this posture, the Earth has per-

performed its mean motion above a revolution requisite to compleat a mean day. For the arch df or the angle $d B f$ is equal to the angle $A C B$ the mean *diurnal* motion of the Earth. Also the arch de , which the Earth, or any *meridian* therein, must pass more than a revolution before it be apparent Noon, is equal to the angle $A S B$ the apparent motion of the Earth at the Sun. From whence 'tis evident, that the arch ef , which the circumference of the rouling Earth performs between the apparent and mean Noon, and which shews the difference between the apparent and mean Day, is equal to the angle $S B C$, which is the *Equation of the Orbit*. Wherefore the *Prosthaphæreses* of the Orbit resolved into parts of time, shall be the *Equations of time*; which *Equations*, throughout this *semicircle of Anomoly* are negative, or to be subtracted from the *apparent time*, for herein the *mean Noon* succeeds the *apparent*.

In like manner, if we take the opposite parts of the Scheme, and consider the Earth in its *Perihelion*. The point g , or the *meridian* $n g$, being made parallel to its yesterdays posture, 'tis plain that the Earth has performed one compleat revolution. This point being carried to h , where 'tis opposite to the center of the Orbit, 'tis now mean Noon; for the arch gh , or the angle $g N h$ equal to the mean *diurnal* motion of the Earth, is passed over. But it is not yet apparent Noon, till the
Earth

Earth by its rouling brings the same *meridian* to *k*, where 'tis directly opposite to the Sun. From whence 'tis manifest that the *apparent Day* exceeds the *mean* by so much time as is requisite for the earth to pass the arch *hk*, which arch is equal to the angle *CNS* the *Prosthapheresis* of the Orbit: wherefore resolving this into time, we have the *Æquation* of time, which throughout this *semicircle* of *Anomaly* is *affirmative*, or to be added to the *apparent* time, because herein the *mean Noon* precedes the *apparent*.

'Tis manifest from what foregoes, that if the Sun were in the center of the great Orbit, and the Earths *Axis* were not *inclined* to its path or way, there would be no *Inequality* of time, but the *mean Day* and *apparent* would be equal. Moreover, if there were no *excentricity* of the Sun from the center of the Orbit, but there were the usual inclination of the Earths *Axis* to the Orbit, tho there would no *Inequality* of time arise such as is shewn in the foregoing demonstration; yet there would arise another *Inequality* from the said inclination of the Earths *Axis*, or as the *Ptolomaicks* would express it, from the inclination of the *Ecliptick* to the *Æquator*; the quantity and affections of which *Inequality* is thus shewn by the *Analemma*.

In Fig. 14. *PCF* is a Quadrant of the *Solstitial Colure*, *P* the Pole, *AF* a Radius of the *Æquator*, *CA* a
Radius

*Radius of the Ecliptick, A the Equinoctial point, or the place of the Sun in the beginning of γ at noon on some certain day, \odot the Suns place at noon the day following; through which place striking the arch $P \odot B$ perpendicular to the $\text{\AE}quator$, $A \odot$ will express the diurnal motion of the Sun, and AB its Right ascension, or the arch of the $\text{\AE}quator$ that Culminates with the Sun. Which arch, seeing 'tis one of the sides of a Right-angled Triangle $A \odot B$, cannot be equal to the Hypotenuse, that is, to the Suns motion $A \odot$. Wherefore seeing the revolutions of the $\text{\AE}quator$, and of its equal or like parts, are equable, and performed in equal times, but the Sun in passing equal parts of the *Ecliptick* applies to the meridian with unequal parts of the $\text{\AE}quator$; it necessarily follows that the solar days are unequal. And that the difference between the Suns true place and its Right ascension being converted into time, is the true $\text{\AE}quation$ of time arising from this cause. Which $\text{\AE}quation$, in the first and third Quadrants of the Zodiack is to be subtracted from the apparent time, for in them the longitude of the Sun from the next Equinoctial point passes the meridian sooner than a like arch projected in the $\text{\AE}quator$. But in the 2d and 4th Quadrants of the Zodiack this $\text{\AE}quation$ is to be added to the apparent time to get the mean; for in these the longitude of the Sun from the $\text{\AE}quinox$ passes the meridian later than the like Arch projected in the*

Æquator. For example, let the *longitude* of the Sun from the first point of γ be $\odot A. 0^{\circ}. 59'. 08''$. its *Right ascension*, or the arch of the *Æquator* culminating therewith *AB* will be $0^{\circ}. 54'. 13''$. their difference $4'. 55''$. being converted into time is $00^h 00'. 19''. 40'''$, and by so much is the *apparent* day shorter than the *mean*. This therefore is the *Æquation* of time arising from this cause, and is negative, or to be subtracted from the *apparent* time, to obtain the *mean* time; for the *longitude* of the Sun arrives at the *meridian* sooner than a like arch projected in the *Æquator*.

Here are therefore demonstrated two sorts of *Æquations* of time arising from two different causes, if they are both to be added, or both to be subtracted, their sum is to be added or subtracted; but if one be to be added, and t'other subtracted, their difference according to the nature of the greatest is to be added or subtracted to or from the *apparent* time to get the *mean*.

And thus far I have presumed to borrow from my Learned and Ingenious Friend's Discourse; which is sufficient, I think, to put this matter out of all dispute.

After clearing the *Theory* of this Doctrine, I come next to apply it to practice in regulating curious Time-keepers, which indeed are very often abused for want of the due consideration and right application of this *Æquation* of time. For
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at some time of the year it happens that if our *Watches* or *Oscillating Pendulums* do not differ above a quarter of an hour from the time shew'd by the Sun or Stars, they are false, and need a correction. And the reason of this is plain, for if a *Pendulum-Watch* goes true, it goes *equal*, that is, one 24 hours at any time of the year, is as long as another 24 hours at any other time of the year, and this perpetually and constantly; that is, all *Watches* that go true, measure the *equal* or *mean* time, and consequently ought to differ from the *apparent* time shewn by a *Sun-dial* or other Instrument, as much as is the *Æquation* of time in excess or defect; but the *Æquation* of time is sometimes above a quarter of an hour, therefore so much ought a good-going *Pendulum-watch* to differ sometimes from the Sun, if it be rightly adjusted. But this will be more evident by explaining the *Tables*. These are calculated by the foregoing *Theory*, and will serve very well for these 20 years to come; though it must be confest, that to have them most accurate, these *Tables* ought to be Calculated for every year, as is manifest to those that consider the foregoing *Theory*. Yet I say, these will serve very well for 20 years to come, without any considerable error. Some few seconds error there may be, and he that desires them more exact, may be at the pains of Calculating them himself; the method whereof he may find laid down

in the forementioned Treatise at the end of *Horrox's Works*; or in *Mr. Flamsteed's Doctrine of the Sphere* publish'd in *Sir J. Moore's System of Mathematicks*. We see there are only four days in the year on which the *æquations* of days cease, that is, the *apparent* and *mean* time are then the same, viz. on the 4th. of *April*, *June* the 6th. *Aug.* the 21st. and *Decemb.* 13. If on any of these days we Set a well-regulated *Pendulum-watch* to the *apparent* time shewn by the Sun or Stars, on any day afterwards, it ought to differ from the *apparent* time shewn by the Sun so much as is the *Æquation* of time in the Table. If the *Æquation* is to be subtracted, the *Pendulum* ought to be so much slower than (or behind) the Sun; if the *Æquation* is to be added, the *Pendulum* ought to be so much faster than (or before) the Sun. For upon any day of the year observing the time exactly by Sun or Stars, that time is the *apparent* time, and to gain the *mean* time which ought to be shewn by the Clock, we are to add too or to subtract from the said *apparent* time; the *Æquation* answering to the day of our observation. Suppose for instance, on the 4th of *April* I observe the time by the Sun, here because the *Æquation* ceases, the *apparent* and *mean* time are the same, and therefore I am to set my Clock to the exact and full time, as the Sun or Stars shew it; but if the *Pendulum* go exactly true, and it move to the 4th of *May*, I shall find it 4'. 17" behind the

the Sun, for so much is the *equation* on the 4th of May to be substracted from the *apparent* time of the Sun to gain the *mean* time of the Clock; that is, when the *Sun* shews it to be 9 a Clock in the morning, the Clock ought to be but 8^h. 55'. 43". And if I find the *Pendulum* more or less behind the *Sun*, it has not gone truly as it ought, but the *Pendulum* or *swagg* is to be lengthned or shortned as is requisite to make it gain or lose the difference betwixt the time shewn by the Clock, and 8^h. 55'. 43". in 30 days elapsed between the 4th of April, and 4th of May; according to a Table, whose use I shall declare presently. But if the Movement be exquisitely true, if it go to the 6th of June, it will again shew the same time with the *Sun* or *Stars*, for then again the *Equation* is nothing. And if it go onwards exactly to the 3^d of August, it will be 4'. 18". before the *Sun*; for at that time so much is the *equation* to be added to the *apparent* time to make it the *mean*. 'Till again on the 24th of October the Watch ought to be 16'. 4". behind the *Sun*, for so much is the *equation* on that day to be substracted.

Wherefore if at any time we set our *Pendulum-Watch* in order to rectify it, and bring it exactly to measure the *mean* day, we are to add to or substract from the *apparent* time shewn by the *Sun* so much as is the *equation* of days at the time we set it. For example, at noon, or just when
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the *Sun* is in the *Meridian* on the 9th of *September*, that is, when the *apparent* time is exactly 12 a Clock, I set my Watch, the *aquation* is then 6'. 26". *subtr.* Wherefore I set my Watch to 11 h. 53'. 34". Which, if it go right, that is *equally* as it ought, on the 9th of *October* will be 14'. 52". behind the the *Sun*; if it be either more or less behind or before the *Sun*, it has gone false, and is to be rectified by lengthning or shortning the *Pendulum* as much as is requisite to make it gain or lose the difference between 14'. 52" behind the *Sun* and its error whatever it is in 30 days time elapsed between the 9th of *Septemb.* and 9th of *Octob.* But if at any other time of the year we set our Watch when the *aquation* is to be added, we must put it so much before the *Sun* as is the *aquation*. But this is plain enough without further Illustration.

Of the certainty and exactness of this *aquation* of time, I have made a most convincing Experiment by an exquisitely rectified *Pendulum-Clock*, which I bought from Mr. Richard Jarrat Watchmaker in *Lothbury, London*, whom I can therefore recommend for his honesty and ability.

And because I have spoken in this Chapter of lengthning or shortning of a *Pendulum*, so as to make it go slower or faster so much in a certain time; for doing this more regularly, and not by guess, I have here added a Table adapted to a *Pendulum*

dulum that *Vibrates seconds*, which is supposed to be 39.2 inches long. Sir *Jonas Moore* in his *Mathematical Compendium*, pag. 113 gives us such a Table as this, but whether by the fault of the Printer or Calculator 'tis very erronious, as any one may find that will be at the pains to examine it by the following Rules for Calculating these Tables. The Rule is, the lengths of *Pendulums* are to each other reciprocally as the squares of their vibrations in the same time. Thus, if a *Pendulum* 39.2 inches long *vibrate* 60 times in a minute, how oft will a *Pendulum* 9.8 (*viz.* quarter of 39.2) inches *vibrate* in a minute? by the foregoing Rule the proportion stands thus, 9.8 : 39.2 :: 3600 : 14400, whole square Root is 120; therefore a *Pendulum* 9.8 will *vibrate* 120 times in a minute. So if it be required how oft a *Pendulum* 39.0 inches *vibrates* in a minute, the *Analogy* will be this, 39.0 : 39.2 :: 3600 : 3618, whose square Root is 60.15. that is, a *Pendulum* 39.0 inches long *vibrates* in a minute 60 times, and 15 hundreds of a vibration more than 60 times. So that multiplying 15 hundreds of a vibration by 1440 the number of minutes in 24 hours, we get the number of vibrations which a *Pendulum* 39.0 inches long *vibrates* in a day more than one of 39.2; and seeing each vibration of the *Pendulum* in a Clock adapted for it, sets the hand forward a second, by knowing the number of vibrations which a *Pendulum* 39.0 inches long performs in a day

day more than a *Pendulum* 39.2 inches long, we may know the number of seconds which it will advance the *Index* of the Clock forward more than one 39.2 inches long. And by these Rules are the following Tables Calculated.

1	2	3	4
38.760".3863	9'. 16". 16'''	1'. 52" 06'''	
38.860.3084	7. 24 . 10	1. 51. 40	
38.960.2309	5. 32 . 30	1. 51. 20	
39.060.1536	3. 41 . 10	1. 50. 40	
39.160.0766	1. 50 . 20	1. 50. 20	
39.260.0000	0. 00 . 00		
39.359.9236	1. 50 . 00	1. 50. 00	
39.459.8475	3. 39 . 36	1. 49. 36	
39.559.7717	5. 28 . 45	1. 49. 09	
39.659.6962	7. 17 . 30	1. 48. 45	
39.759.6210	9. 05 . 45	1. 48. 15	

The first Column has in the middle the length of the *Pendulum* 39.2 ; upwards it diminisheth one tenth, and downwards it increaseth one tenth. The second Column is the *vibrations* and parts of a *Vibration* performed in a minute by the lengths in

in the first. The third Column is the minutes and seconds that these lengthnings and shortnings of the *Pendulum* will cause in a day, and are gotten by multiplying 1440' the minutes in a day by the *Decimals* above or under 60". The 4th and last Column are the differences of the 3^d. The like Table may be made to any length of a *Pendulum*, respect being had to the foregoing Rule.

I shall conclude all relating to my *Dial* with the Calculation of hours and minutes for an *Horizontal Dial* for the *latitude* of *Dublin* $53^{\circ}.20'$. Which I believe will not be unacceptable to those that design curious *Dials* for that place.

The Tables follow amongst the others.

CH A P. XI.

Of the Tables of the Circumpolar Stars, their Calculation and Uses.

I Shall here, as a conclusion to this Work, add something concerning very useful Tables for shewing the time of night very accurately, and other operations by the *Circumpolars*, or Stars that never set in our *latitude* of *Dublin*. Such Tables as these we have mentioned in Sir J. Moore's *Mathematical Compendium* p. 118, 119. but they are for the *latitude* of *London*; neither does he give us all their

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uses, or the method of Calculating them for other places. I shall do both in this place, and first for their use.

In any Northern Window, or other convenient place, hang up a good weighty Plummert by a fine and even silk or silver-wire, then placing your eye at some distance behind this thred, that is, to the South thereof, observe when any of the Stars mentioned in the Table are cut by this thred at the same time with the *Stole-Star*. From the Stars Right ascension in the Table (adding 24 hours if need be) subtract the Suns right ascension, the remainder gives the hour, minute, and second of the night. The Tables consist of 6 Columns, the first shews the *Magnitudes*; the second contains the Stars names, as they are described on our common *English Globes*; the third shews the *Right ascension* of the *Mid-heaven* in time, when any of these Stars come in the same *Vertical* with the *Pole-Star*, by which (working according to the foregoing Rule) we find the hour of the night; the fourth Column gives the difference in time between the coming of any of these Stars under the *Pole-star*, and their coming under the *Pole* it self. Some of these Stars pass the *Meridian*, or come under the *Pole* before they come under the *Pole-Star*, such as are all the Stars whose *Right ascensions* are above $9^{\circ}. 14'. 10''$, and under $189^{\circ}. 14'. 10''$. And of these we say nothing in this place, only they are all marked in the Table

e with *E*, as having their *Azimuth* when under
 e *Pole-star Eastward*. But of these Stars in this
 catalogue which pass the *Meridian*, or come under
 e *Pole* after they have left the *Vertical* of the
 le-star, we make this following use. All these
 e marked in the sixth Column with *W*, as ha-
 ng a *Western Azimuth* when they are under the
 le-star. Wherefore placing a second line and
 mmet behind, or to the *South* of the former, then
 serve by the first when any of these Stars come
 der the *Pole-star*, and by a *Pendulum* from that
 stant count the time in this fourth Column,
 d placing your eye now behind both the lines,
 st at the end of your count by gently moving
 is last thred, make them both cut the *Star* you
 unt for, then are these two lines exactly in the
 ridian; and is a curious way for finding a *Me-*
 lian line. The fifth and sixth Columns serve for
 e same use; for when any of these *Stars* are un-
 r the *Pole-star*, making your two lines cut both
 le-star and t'other, these two lines hang so far
 ut of the *Meridian line*, as is the *Azimuth* expressed
 the Table, which *Azimuth* is shewn to be *East*
 West by the 6th Column. Wherefore this angle
 eing set off from the found line, shews the true
Meridian. And we may observe, that there are
 three stars expressed in the Table, which being
 under the *pole-star* are insensibly nigh the *Meri-*
 dian, these are *Cassiopeia's Hipp*, *Cor Caroli*, and *Aliot*

uses, or the method of Calculating them for other places. I shall do both in this place, and first for their use.

In any Northern Window, or other convenient place, hang up a good weighty Plummert by fine and even silk or silver-wire, then placing your eye at some distance behind this thred, that is, to the South thereof, observe when any of the Stars mentioned in the Table are cut by this thred at the same time with the *Stole-Star*. From the *Star* Right ascension in the Table (adding 24 hours if need be) subtract the *Sun's* right ascension, the remainder gives the hour, minute, and second of the night. The Table consist of 6 Columns, the first shews the *Magnitudes*; the second contains the *Stars* names, as they are described on our common *English Globes*; the third shews the *Right ascension* of the *Mid-heaven* in time, when any of these *Stars* come in the same *Vertical* with the *Pole-Star*, by which (working according to the foregoing Rule) we find the hour of the night; the fourth Column gives the difference in time between the coming of any of these *Stars* under the *Pole-star*, and their coming under the *Pole* it self. Some of these *Stars* pass the *Meridian*, or come under the *Pole* before they come under the *Pole-Star*, such as are all the *Stars* whose *Right ascensions* are above $9^{\circ}. 14'. 10''$, and under $189^{\circ}. 14'. 10''$. And of these we say nothing in this place, only they are all marked in the Table

ble with *E*, as having their *Azimuth* when under the *Pole-star Eastward*. But of these Stars in this Catalogue which pass the *Meridian*, or come under the *Pole* after they have left the *Vertical* of the *Pole-star*, we make this following use. All these are marked in the sixth Column with *W*, as having a *Western Azimuth* when they are under the *Pole-star*. Wherefore placing a second line and plummet behind, or to the *South* of the former, then observe by the first when any of these Stars come under the *Pole-star*, and by a *Pendulum* from that instant count the time in this fourth Column, and placing your eye now behind both the lines, just at the end of your count by gently moving this last thred, make them both cut the *Star* you count for, then are these two lines exactly in the *meridian*; and is a curious way for finding a *Meridian line*. The fifth and sixth Columns serve for the same use; for when any of these Stars are under the *Pole-star*, making your two lines cut both *Pole-star* and t'other, these two lines hang so far out of the *Meridian line*, as is the *Azimuth* expressed in the Table, which *Azimuth* is shewn to be *East* or *West* by the 6th Column. Wherefore this angle being set off from the found line, shews the true *Meridian*. And we may observe, that there are three stars expressed in the Table, which being under the *pole-star* are insensibly nigh the *Meridian*, these are *Cassiopeia's Hipp*, *Cor Caroli*, and *Aliot*

or the *Great Bears Rump*. Wherefore whenever we have any of these under the *pole-star*, we may by it find a *Meridian line* without any sensible error. Another use that we may make of this Table, is for the true adjusting of *Hour-glasses* and other *Time-keepers*, as *Pendulum-Watches*, &c. for trying their going, and bringing them to their right measure. Thus we shall find it just four hours (wanting one second) between the coming of *Cor Caroli*, and the 24th Star of *Draco* (in *Tich. Catal.*) under the *Pole-star*; for when *Cor Caroli* is under the *pole-star*, the *Right Ascension* of the *Mid-heaven* in time is 00h. 36'. 44". And when the 24th of *Draco* is under the *pole-star*, the *mid-heavens temporary Right Ascension* is 4h. 36'. 43". their difference is 3h. 59'. 59". which wants only one second of four hours. Note that this 24th of *Draco* is marked in *Bayer* by α .

I come now to shew the method of Calculating this Table, for which see the 15 and 16 Figures. Wherein $axbp$ is the circle described by the *pole-star* x round the *pole* p , z the *Zenith*, zn the *meridian*, whose *North* part is pn , s is any other *Circumpolar stars*, zxs a *Vertical circle*, hn the *Horizon*, eq the *Æquator*, po the *Axis Mundi*. There are here given (or at least may be known from the Tables) px the Complement of the *pole-stars Declination*, ps the Complement of the other *stars Declination*, xps the difference of the *Right Ascensions*
of

of the *pole-star*, and other *star*, also zp the Complement of the *Poles Elevation*. Wherefore, in the *Triangle* spx , having sp , px , and the angle spx , we may find the angle sxp ; and having that, we have the angle pxz , for this is the Complement of sxp to 180 degrees. Then in the *Triangle* zpx , we have zp , px , and the angle pxz , to find the angle pzx , which is the *Azimuth* that these two *Stars* have, when they come in the same *Vertical*, and makes the fifth Column in our Table. And from the same *Data*, we may find the angle zpx . Therefore in the *Triangle* zps , we have found the angle zps , for zps is equal to the sum of zpx and spx . Now the difference between zps and 180 degrees is equal to the difference in time between the *Star* s coming under the *Pole-star* and coming, under the *Pole*, or its being in the *Meridian*, and consequently its having its own *Right Ascension*, or its Complement; and this makes the fourth Column in our Table. Now this difference in time between the *Pole* and *Pole-star* being added to or subtracted from the true *Right Ascension* of the *Star* s in time, gives the *Right Ascension* in time of the *Star* s (or of the *mid-heaven*) when 'tis under the *Pole-star*; and this makes our third Column in the Table. The Rule to know when this difference in time between the *Pole* and *Pole-star* is to be added, and when to be subtracted, will be evident by observing, that all such *Stars* whose *Right Ascen-*

Ascensions are above the *Right Ascension* of the *Pole-star* (for the year, as for 1680) $9^{\circ}. 14'. 10''$, and under $189^{\circ}. 14'. 10''$. pass the *meridian* before they come under the *Pole-star*; all the other *semicircle*, contrarily; for if the *Star* pass the *meridian* before it comes under the *Pole-star*, this difference in time between the *Pole* and *Pole-star* is to be added to the *Right Ascension* of the *Star*; if the contrary, 'tis to be substracted.

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Ascensions are above the *Right Ascension* of the *Pole-star* (for the year, as for 1680) $9^{\circ}. 14'. 10''$, and under $189^{\circ}. 14'. 10''$. pass the *meridian* before they come under the *Pole-star*; all the other *semicircle*, contrarily; for if the *Star* pass the *meridian* before it comes under the *Pole-star*, this difference in time between the *Pole* and *Pole-star* is to be added to the *Right Ascension* of the *Star*; if the contrary, 'tis to be subtracted.

F I N I S.

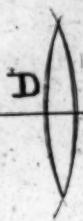


Fig. 10

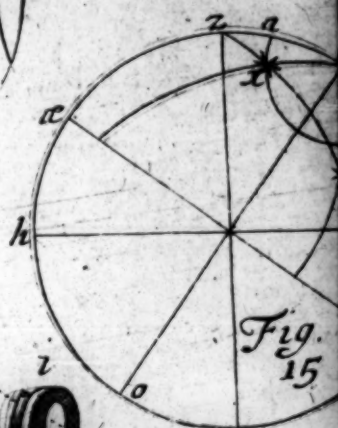


Fig. 15

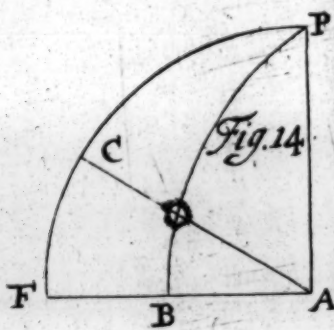


Fig. 14

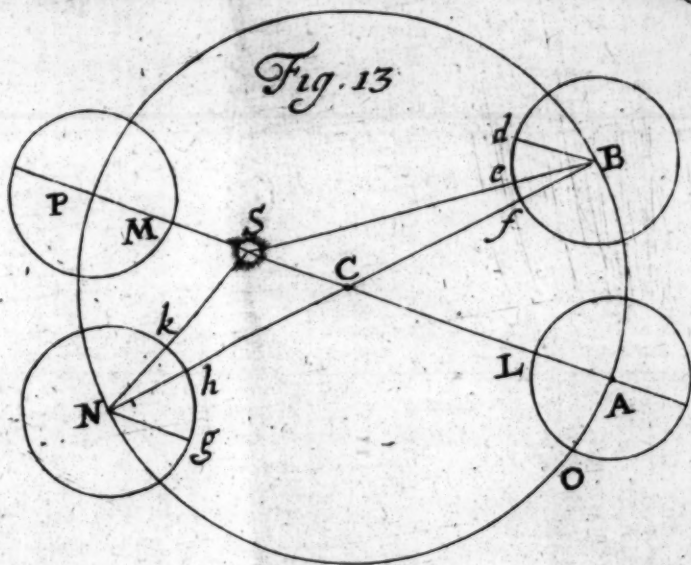


Fig. 13

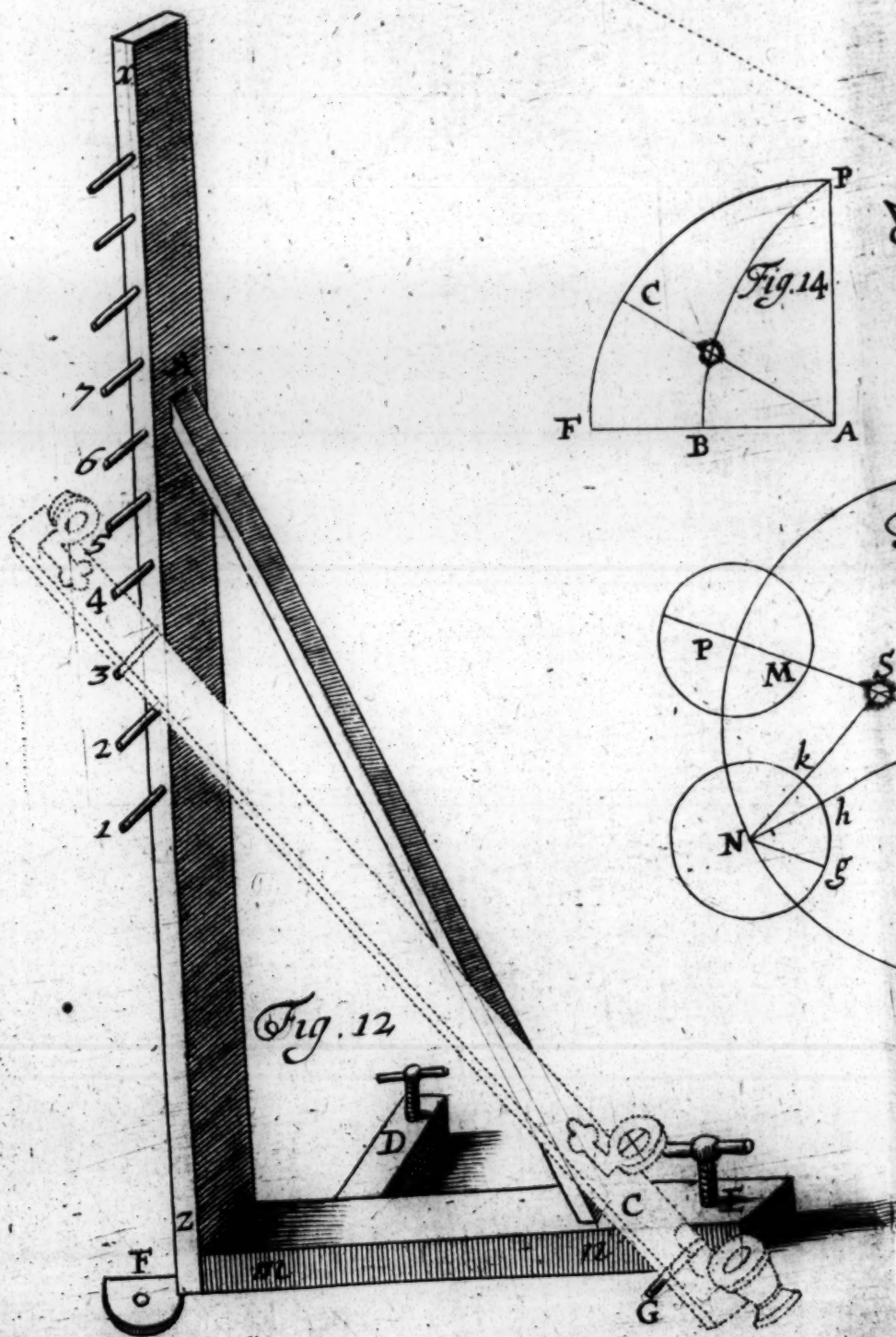


Fig. 12

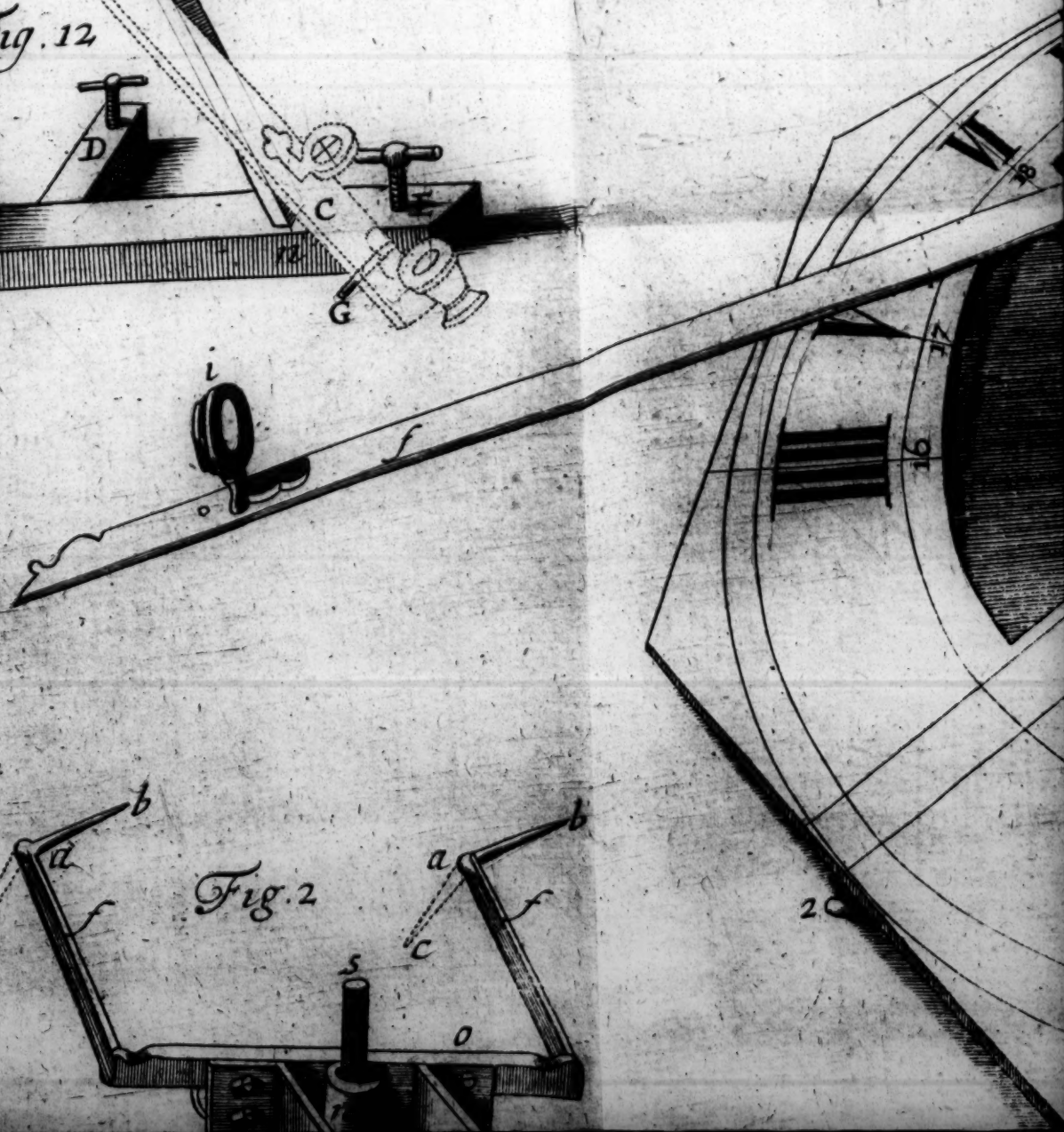


Fig. 2

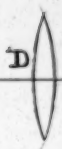
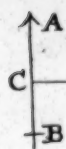


Fig. 10

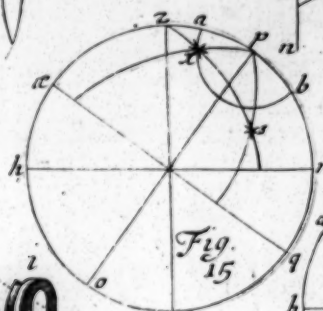
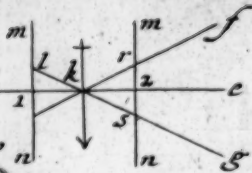


Fig. 15

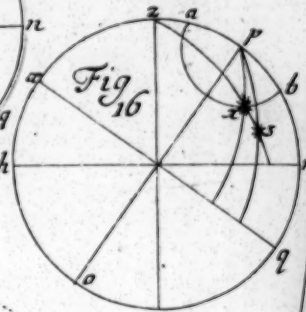


Fig. 16

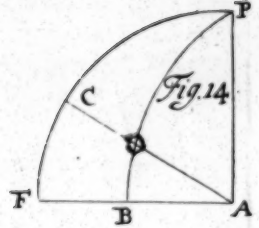


Fig. 14

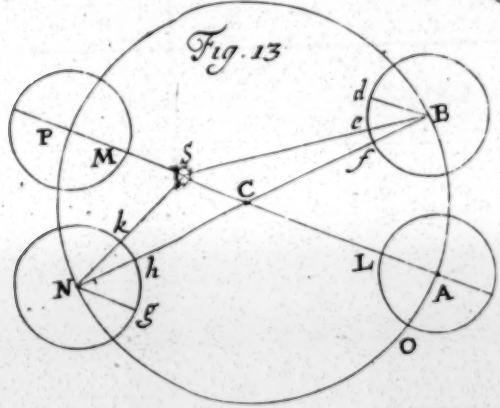


Fig. 13

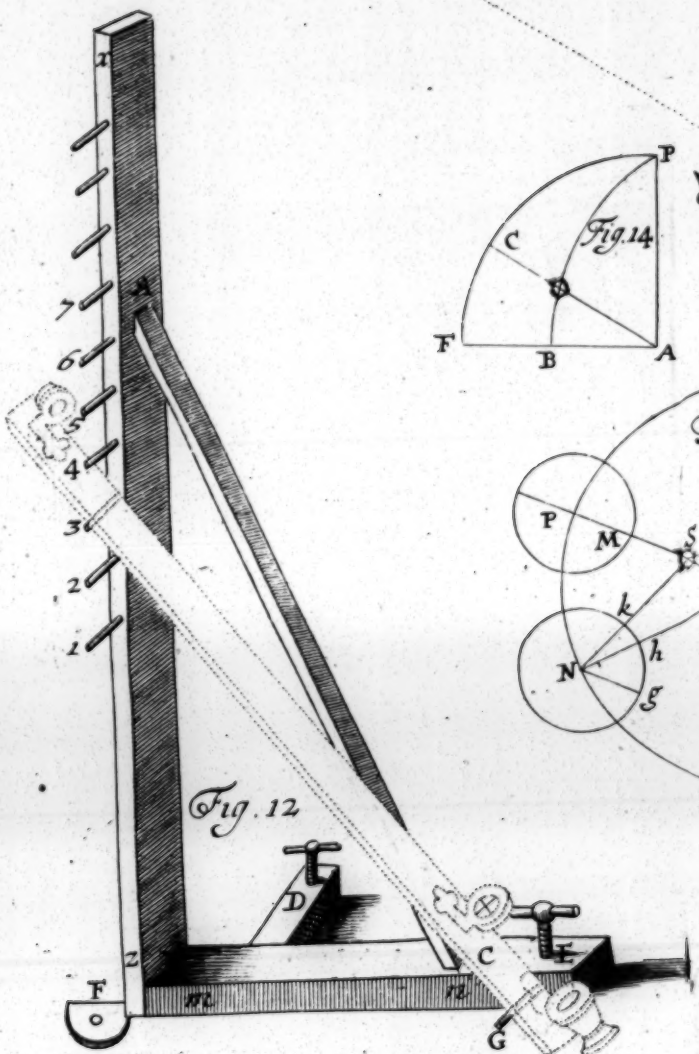


Fig. 12

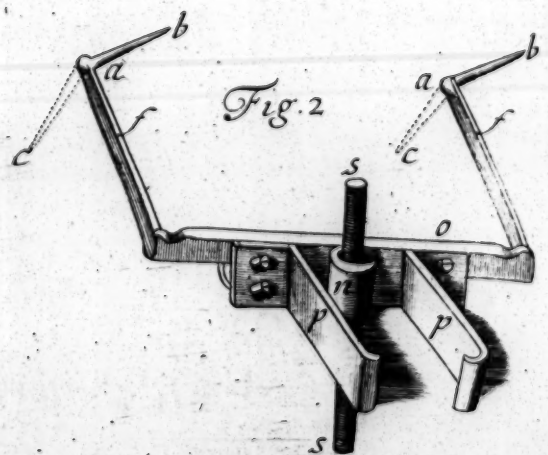
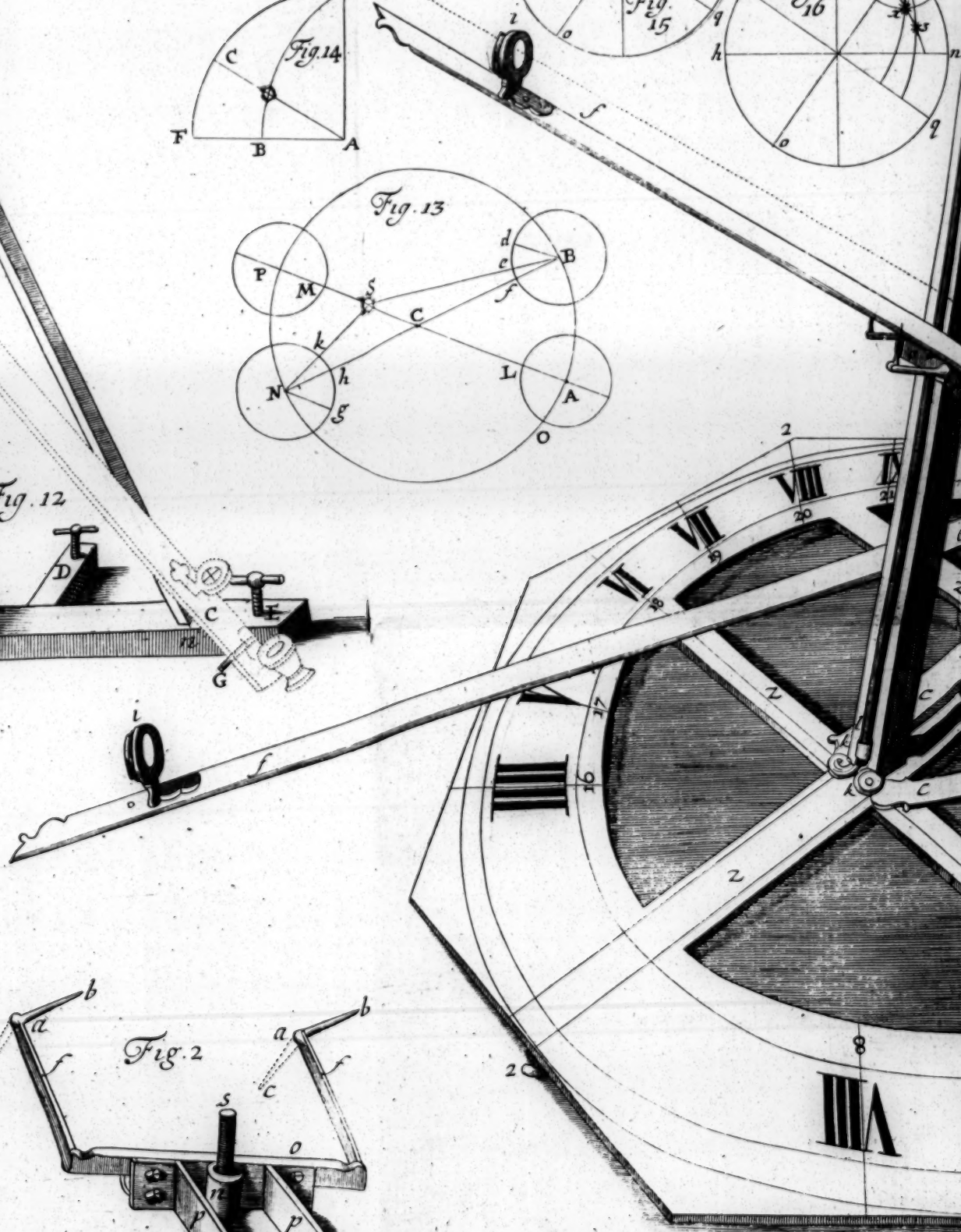
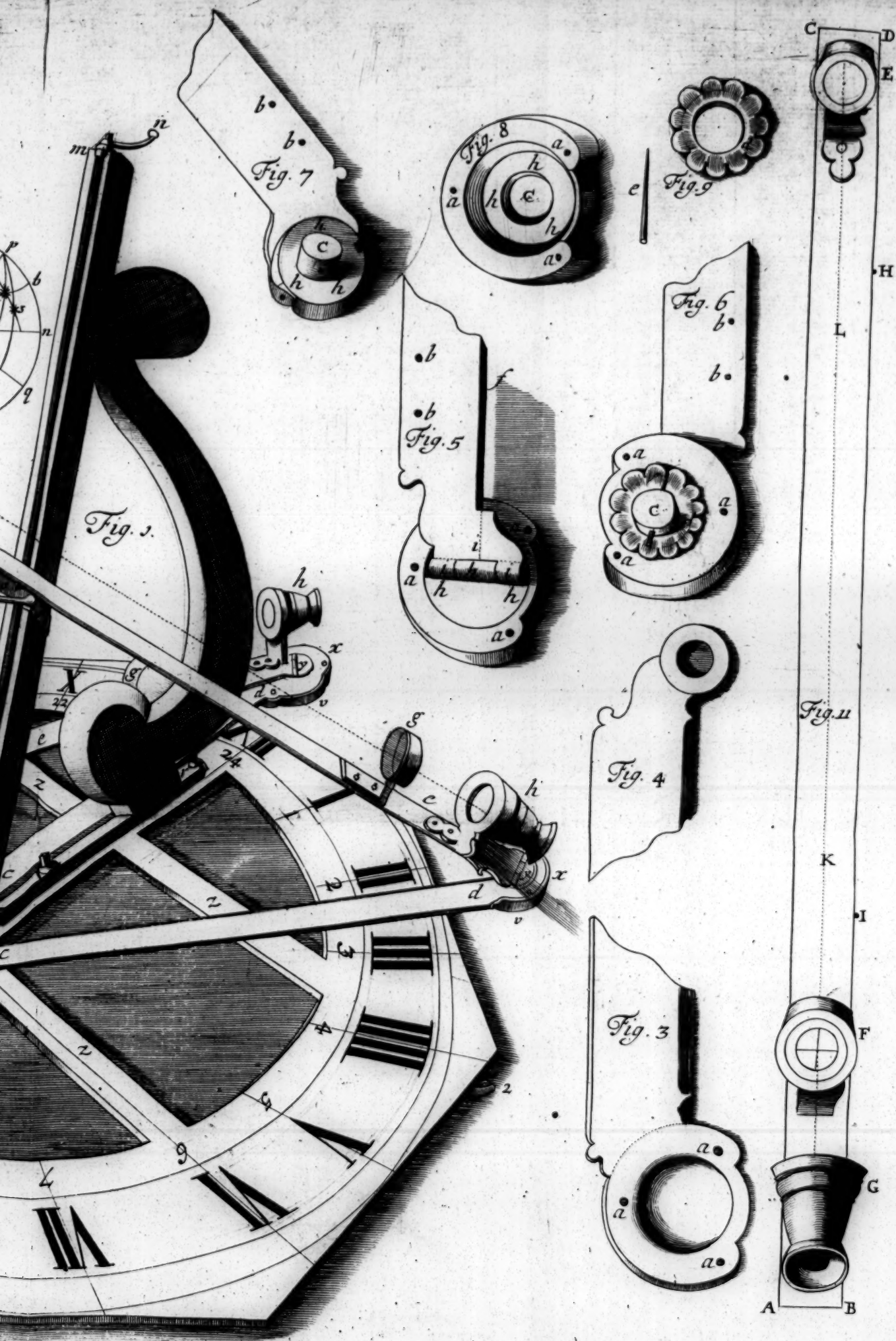


Fig. 2



To be inserted at the End of the book before the Tables.

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T A B L E S
OF THE
Suns Right Ascension in Time
TO EVERY
Ten Minutes of the ECLIPTICK.

TABLES

OF THE

Sun's Right Ascension in Time

TO EVERY

Ten Minutes of the ECLIPSE.

D. m.	γ H	μ H	Com. Pn. " "	X
0. 0	0	12	0. 0	
10	0	12	0.37	"
20	0	12	1.13	37
30	0	12	1.50	
40	0	12	2.27	
50	0	12	3. 3	
1. 0	0	12	3.40	
10	0	12	4.17	"
20	0	12	4.53	37
30	0	12	5.30	
40	0	12	6. 7	
50	0	12	6.43	
2. 0	0	12	7.20	
10	0	12	7.57	"
20	0	12	8.33	37
30	0	12	9.10	
40	0	12	9.46	
50	0	12	10.23	
3. 0	0	12	11. 0	
10	0	12	11.37	"
20	0	12	12.13	37
30	0	12	12.50	
40	0	12	13.26	
50	0	12	14. 3	
4. 0	0	12	14.40	
10	0	12	15.17	"
20	0	12	15.53	37
30	0	12	16.30	
40	0	12	17. 6	
50	0	12	17.43	

D. m.	γ H	μ H	Com. Pn. " "	X
5. 0	0	12	18.20	
10	0	12	18.57	"
20	0	12	19.33	37
30	0	12	20.10	
40	0	12	20.46	
50	0	12	21.23	
6. 0	0	12	22. 0	
10	0	12	22.37	"
20	0	12	23.13	37
30	0	12	23.50	
40	0	12	24.26	
50	0	12	25. 3	
7. 0	0	12	25.41	
10	0	12	26.18	"
20	0	12	26.55	37
30	0	12	27.32	
40	0	12	28. 9	
50	0	12	28.46	
8. 0	0	12	29.23	
10	0	12	30. 0	"
20	0	12	30.36	37
30	0	12	31.13	
40	0	12	31.50	
50	0	12	32.26	
9. 0	0	12	33. 3	
10	0	12	33.40	"
20	0	12	34.16	37
30	0	12	34.53	
40	0	12	35.30	
50	0	12	36. 7	

D. m.	γ H	α H	Com. Pts. ' "	X
10. 0	0	12	36.44	" 37
10	0	12	37.21	
20	0	12	37.58	
30	0	12	38.35	
40	0	12	39.12	
50	0	12	39.49	
11. 0	0	12	40.26	" 37
10	0	12	41. .3	
20	0	12	41.40	
30	0	12	42.17	
40	0	12	42.54	
50	0	12	43.31	
12. 0	0	12	44. 7	" 37
10	0	12	44.44	
20	0	12	45.21	
30	0	12	45.58	
40	0	12	46.35	
50	0	12	47.12	
13. 0	0	12	47.49	" 37
10	0	12	48.26	
20	0	12	49.03	
30	0	12	49.40	
40	0	12	50.17	
50	0	12	50.54	
14. 0	0	12	51.31	" 37
10	0	12	52. 8	
20	0	12	52.45	
30	0	12	53.22	
40	0	12	53.59	
50	0	12	54.36	

D. m.	γ H	α H	Com. Pts. ' "	X
15. 0	0	12	55.13	" 37
10	0	12	55.50	
20	0	12	56.27	
30	0	12	57. 4	
40	0	12	57.41	
50	0	12	58.19	
16. 0	0	12	58.5	" 37
10	0	12	59.3	
20	1	13	00. 0	
30	1	13	00.47	
40	1	13	1.24	
50	1	13	2. 2	
17. 0	1	13	2.39	" 37
10	1	13	3.16	
20	1	13	3.53	
30	1	13	4.30	
40	1	13	5. 7	
50	1	13	5.45	
18. 0	1	13	6.22	" 37
10	1	13	6.59	
20	1	13	7.37	
30	1	13	8.14	
40	1	13	8.52	
50	1	13	9.29	
19. 0	1	13	10. 6	" 37
10	1	13	10.43	
20	1	13	11.21	
30	1	13	11.58	
40	1	13	12.36	
50	1	13	13.13	

D. m.	V	H	Com. Pts.	
20.0	I	13	13.50	
10	I	13	14.27	"
20	I	13	15.5	38
30	I	13	15.42	
40	I	13	16.20	
50	I	13	16.57	
21.0	I	13	17.34	
10	I	13	18.12	"
20	I	13	18.50	38
30	I	13	19.27	
40	I	13	20.5	
50	I	13	20.43	
22.0	I	13	21.19	
10	I	13	21.57	"
20	I	13	22.35	39
30	I	13	23.12	
40	I	13	23.50	
50	I	13	24.28	
23.0	I	13	25.5	
10	I	13	25.42	"
20	I	13	26.20	39
30	I	13	26.57	
40	I	13	27.35	
50	I	13	28.13	
24.0	I	13	28.50	
10	I	13	29.28	"
20	I	13	30.6	39
30	I	13	30.44	
40	I	13	31.22	
50	I	13	32.0	

D. m.	V	H	Com. Pts.	
25.0	I	13	32.37	
10	I	13	33.15	"
20	I	13	33.53	38
30	I	13	34.31	
40	I	13	35.9	
50	I	13	35.46	
26.0	I	13	36.24	
10	I	13	37.2	"
20	I	13	37.40	38
30	I	13	38.18	
40	I	13	38.56	
50	I	13	39.33	
27.0	I	13	40.11	
10	I	13	40.49	"
20	I	13	41.27	38
30	I	13	42.5	
40	I	13	42.43	
50	I	13	43.20	
28.0	I	13	43.99	
10	I	13	44.37	"
20	I	13	45.15	38
30	I	13	45.53	
40	I	13	46.31	
50	I	13	47.9	
29.0	I	13	47.47	
10	I	13	48.25	"
20	I	13	49.03	39
30	I	13	50.41	
40	I	13	51.19	
50	I	13	51.58	

D. m.	8 H	M H	Com. Pts. 1 "
0. 0	1	13	51.36
10	1	13	52.14
20	1	13	52.52
30	1	13	53.30
40	1	13	54. 8
50	1	13	54.47
1. 0	1	13	55.25
10	1	13	56. 4
20	1	13	56.42
30	1	13	57.21
40	1	13	57.59
50	1	13	58.38
2. 0	1	13	59.15
10	1	13	59.54
20	2	14	0.32
30	2	14	1.11
40	2	14	1.49
50	2	14	2.27
3. 0	2	14	3. 6
10	2	14	3.45
20	2	14	4.23
30	2	14	5. 1
40	2	14	5.39
50	2	14	6.18
4. 0	2	14	6.57
10	2	14	7.36
20	2	14	8.15
30	2	14	8.53
40	2	14	9.32
50	2	14	10.11

D. m.	8 H	M H	Com. Pts. 1 "
5. 0	2	14	10.49
10	2	14	11.28
20	2	14	12. 7
30	2	14	12.46
40	2	14	13.25
50	2	14	14. 4
6. 0	2	14	14.42
10	2	14	15.21
20	2	14	16. 0
30	2	14	16.39
40	2	14	17.18
50	2	14	17.57
7. 0	2	14	18.35
10	2	14	19.14
20	2	14	19.53
30	2	14	20.32
40	2	14	21.11
50	2	14	21.50
8. 0	2	14	22.29
10	2	14	23. 8
20	2	14	23.47
30	2	14	24.26
40	2	14	25. 5
50	2	14	25.44
9. 0	2	14	26.24
10	2	14	27. 3
20	2	14	27.41
30	2	14	28.20
40	2	14	28.59
50	2	14	29.38

D. m.	8 H	11 H	Com. Pts. ! "
10.0	2	14	30.19
10	2	14	30.59 "
20	2	14	31.38 39
30	2	14	32.17
40	2	14	33.7
50	2	14	33.46
11.0	2	14	34.15
10	2	14	34.54 "
20	2	14	35.34 39
30	2	14	36.13
40	2	14	36.51
50	2	14	37.31
12.0	2	14	38.11
10	2	14	38.51 "
20	2	14	39.31 40
30	2	14	40.11
40	2	14	40.51
50	2	14	41.31
13.0	2	14	42.9
10	2	14	42.49 "
20	2	14	43.29 40
30	2	14	44.8
40	2	14	44.48
50	2	14	45.27
14.0	2	14	46.7
10	2	14	46.47 "
20	2	14	47.27 40
30	2	14	48.6
40	2	14	48.46
50	2	14	49.25

D m.	8 H	11 H	Com. pts. ! "
15.0	2	14	50.5
10	2	14	50.45 "
20	2	14	51.25 40
30	2	14	52.5
40	2	14	52.45
50	2	14	53.25
16.0	2	14	54.5
10	2	14	54.45 "
20	2	14	55.25 40
30	2	14	56.5
40	2	14	56.45
50	2	14	57.25
17.0	2	14	58.5
10	2	14	58.45 "
20	2	14	59.25 40
30	3	15	0.5
40	3	15	0.46
50	3	15	1.26
18.0	3	15	2.6
10	3	15	2.46 "
20	3	15	3.27 40
30	3	15	4.7
40	3	15	4.47
50	3	15	5.28
19.0	3	15	6.8
10	3	15	6.49 "
20	3	15	7.29 40
30	3	15	8.10
40	3	15	8.50
50	3	15	9.30

D. m.	δ H	μ H	Com. ps. I''
20.0	3	15	10.10
10	3	15	10.51
20	3	15	11.30
30	3	15	12.11
40	3	15	12.51
50	3	15	13.32
21.0	3	15	14.13
10	3	15	14.53
20	3	15	15.34
30	3	15	16.15
40	3	15	16.55
50	3	15	17.36
22.0	3	15	18.17
10	3	15	18.58
20	3	15	19.39
30	3	15	20.20
40	3	15	21.1
50	3	15	21.41
23.0	3	15	22.22
10	3	15	23.03
20	3	15	23.44
30	3	15	24.25
40	3	15	25.06
50	3	15	25.47
24.0	3	15	26.27
10	3	15	27.08
20	5	15	27.49
30	3	15	28.30
40	3	15	29.11
50	3	15	29.52

D. m.	δ H	μ H	Com. ps. I''
25.0	3	15	30.33
10	3	15	31.14
20	3	15	31.55
30	3	15	32.36
40	3	15	33.17
50	3	15	33.58
26.0	3	15	34.40
10	3	15	35.21
20	3	15	36.02
30	3	15	36.44
40	3	15	37.25
50	3	15	38.06
27.0	3	15	38.47
10	3	15	39.29
20	3	15	40.11
30	3	15	40.53
40	3	15	41.34
50	3	15	42.15
28.0	3	15	42.55
10	3	15	43.37
20	3	15	44.19
30	3	15	45.01
40	3	15	45.43
50	3	15	46.21
29.0	3	15	47.4
10	3	15	47.46
20	3	15	48.28
30	3	15	49.10
40	3	15	49.52
50	3	15	50.33

D. m.	II H	I H	Com. Fra. P. II
0	3	15	51.14
10	3	15	51.56
20	3	15	52.37
30	3	15	53.19
40	3	15	54.0
50	3	15	54.42
1. 0	3	15	55.24
10	3	15	56.6
20	3	15	56.47
30	3	15	57.29
40	3	15	58.10
50	3	15	58.52
2. 0	3	15	59.35
10	4	16	00.17
20	4	16	00.59
30	4	16	01.40
40	4	16	2.22
50	4	16	3.4
3. 0	4	16	3.46
10	4	16	4.28
20	4	16	5.9
30	4	16	5.51
40	4	16	6.33
50	4	16	7.15
4. 0	4	16	7.59
10	4	16	8.41
20	4	16	9.23
30	4	16	10.5
40	4	16	10.47
50	4	16	11.29

D. m.	II H	I H	Com. pts. P. II
5. 0	4	16	12.11
10	4	16	12.53
20	4	16	13.35
30	4	16	14.17
40	4	16	14.59
50	4	16	15.41
6. 0	4	16	16.25
10	4	16	17.7
20	4	16	17.50
30	4	16	18.32
40	4	16	19.15
50	4	16	19.57
7. 0	4	16	20.39
10	4	16	21.21
20	4	16	22.4
30	4	16	22.47
40	4	19	23.30
50	4	16	24.12
8. 0	4	16	24.54
10	4	16	25.36
20	4	16	26.19
30	4	16	27.1
40	4	16	27.44
50	4	16	28.26
9. 0	4	16	29.9
10	4	16	29.51
20	4	16	30.34
30	4	16	31.16
40	4	16	31.59
50	4	16	32.42

D. m.	H	H	Com. per " "
10.0	4	16	33.25
10	4	16	34.8 "
20	4	16	34.51 43
30	4	16	35.34
40	4	16	36.15
50	4	16	36.58
11.0	4	16	37.41
10	4	16	38.24 "
20	4	16	39.7 43
30	4	16	39.49
40	4	16	40.32
50	4	16	41.14
12.0	4	16	41.57
10	4	16	42.40 "
20	4	16	43.23 43
30	4	16	44.05
40	4	16	44.48
50	4	16	45.31
13.0	4	16	46.15
10	4	16	46.58 "
20	4	16	47.41 43
30	4	16	48.23
40	4	16	49.5
50	4	16	49.48
14.0	4	16	50.32
10	4	16	51.15 "
20	4	16	51.58 43
30	4	16	52.41
40	4	16	53.23
50	4	16	54.5

D. m.	H	H	Com. per " "
15.0	4	16	54.51
10	4	16	55.34 "
20	4	16	56.17 43
30	4	16	57.00
40	4	16	57.43
50	4	16	58.26
16.0	4	16	59.9
10	4	16	59.52 "
20	5	6	00.35 43
30	5	6	1.18
40	5	6	2.01
50	5	6	2.44
17.0	5	17	3.28
10	5	17	4.11 "
20	5	17	4.54 43
30	5	17	5.37
40	5	17	6.20
50	5	17	7.04
18.0	5	17	7.48
10	5	17	8.31 "
20	5	17	9.14 43
30	5	17	9.58
40	5	17	10.41
50	5	17	11.24
19.0	5	17	12.8
10	5	17	12.52 "
20	5	17	13.35 43
30	5	17	14.18
40	5	17	15.01
50	5	17	15.44

D. m.	H	H	Com. Pts.	
20.0	5	17	16.28	
10	5	17	17.12	"
20	5	17	17.55	
30	5	17	18.39	44
40	5	17	19.23	
50	5	17	20.05	
21.0	5	17	20.48	
10	5	17	21.32	"
20	5	17	22.15	
30	5	17	22.59	44
40	5	17	23.42	
50	5	17	24.26	
22.0	5	17	25.09	
10	5	17	25.53	"
20	5	17	26.37	
30	5	17	27.21	44
40	5	17	28.04	
50	5	17	28.47	
23.0	5	17	29.30	
10	5	17	30.14	"
20	5	17	30.58	
30	5	17	31.41	44
40	5	17	32.25	
50	5	17	33.08	
24.0	5	17	33.51	
10	5	17	34.35	"
20	5	17	35.18	
30	5	17	36.02	44
40	5	17	36.46	
50	5	17	37.30	

D. m.	H	H	Com. Pts.	
25.0	5	17	38.12	
10	5	17	38.56	"
20	5	17	39.39	44
30	5	17	40.23	
40	5	17	41.06	
50	5	17	41.50	
26.0	5	17	42.33	
10	5	17	43.17	"
20	5	17	44.00	44
30	5	17	44.44	
40	5	17	45.27	
50	5	17	46.11	
27.0	5	17	46.55	
10	5	17	47.39	"
20	5	17	48.22	44
30	5	17	49.06	
40	5	17	49.49	
50	5	17	50.33	
28.0	5	17	51.17	
10	5	17	52.01	"
20	5	17	52.44	44
30	5	17	53.27	
40	5	17	54.11	
50	5	17	54.54	
29.0	5	17	55.38	
10	5	17	56.22	"
20	5	17	57.06	44
30	5	17	57.49	
40	5	17	58.33	
50	5	17	59.16	

D. m.	S	H	W	Com. Pts.	
0.	0	6	18	0.0	
10	6	18		00.44	"
20	6	18		1.27	44
30	6	18		2.01	
40	6	18		2.44	
50	6	18		3.28	
1.	0	6	18	4.22	
10	6	18		5.06	"
20	6	18		5.50	44
30	6	18		6.34	
40	6	18		7.18	
50	6	18		8.02	
2.	0	6	18	8.43	
10	6	18		9.26	"
20	6	18		10.10	44
30	6	18		10.54	
40	6	18		11.37	
50	6	18		12.20	
3.	0	6	18	13.3	
10	6	18		13.49	"
20	6	18		14.32	44
30	6	18		15.16	
40	6	18		15.59	
50	6	18		16.43	
4.	0	6	18	17.27	
10	6	18		18.11	"
20	6	18		18.54	44
30	6	18		19.38	
40	6	18		20.22	
50	6	18		21.05	

D. m.	S	H	W	Com. Pts.	
5.	0	6	18	21.48	
10	6	18		22.32	"
20	6	18		23.15	44
30	6	18		23.59	
40	6	18		24.42	
50	6	18		25.25	
6.	0	6	18	26.09	
10	6	18		26.53	"
20	6	18		27.37	44
30	6	18		28.20	
40	6	18		29.04	
50	6	18		29.47	
7.	0	6	18	30.30	
10	6	18		31.14	"
20	6	18		31.57	44
30	6	18		32.41	
40	6	18		33.24	
50	6	18		34.17	
8.	0	6	18	34.91	
10	6	18		35.34	"
20	6	18		36.18	44
30	6	18		37.01	
40	6	18		37.45	
50	6	18		38.28	
9.	0	6	18	39.11	
10	6	18		39.54	"
20	6	18		40.48	44
30	6	18		41.31	
40	6	18		42.15	
50	6	18		42.58	

D. m.	S H	V H	Com. Pts. H H	
10. 0	6	18	43.32	
10	6	18	44.15	"
20	6	18	44.59	43
30	6	18	45.42	
40	6	18	46.25	
50	6	18	47. 9	
11. 0	6	18	47.52	
10	6	18	48.35	"
20	6	18	49.18	43
30	6	18	50. 2	
40	6	18	50.45	
50	6	18	51.29	
12. 0	6	18	52.12	
10	6	18	52.55	"
20	6	18	53.38	43
30	6	18	54.22	
40	6	18	55. 5	
50	6	18	55.48	
13. 0	6	18	56.31	
10	6	18	57.14	"
20	6	18	57.57	43
30	6	18	58.41	
40	6	18	59.24	
50	7	19	0 7	
14. 0	7	19	0.50	
10	7	19	1.33	"
20	7	19	2.16	43
30	7	19	3. 0	
40	7	19	3.43	
50	7	19	4.26	

D. m.	S H	V H	Com. Pts. H H	
15. 0	7	19	5. 9	
10	7	19	5.52	"
20	7	19	6.35	43
30	7	19	7.18	
40	7	19	8. 1	
50	7	19	8.44	
16. 0	7	19	9.27	
10	7	19	10.10	"
20	7	19	10.53	43
30	7	19	11.36	
40	7	19	12.19	
50	7	19	13. 2	
17. 0	7	19	13.45	
10	7	19	14.28	"
20	7	19	15.11	43
30	7	19	15.54	
40	7	19	16.36	
50	7	19	17.19	
18. 0	7	19	18. 2	
10	7	19	18.45	"
20	7	19	19.28	43
30	7	19	20.11	
40	7	19	20.53	
50	7	19	21.36	
19. 0	7	19	22.19	
10	7	19	23. 2	"
20	7	19	23.45	43
30	7	19	24.27	
40	7	19	25.10	
50	7	19	25.53	

D. m.	S H	W H	Com. Pa. 1 11
20. 0	7	19	26.35
10	7	19	27.18 //
20	7	19	28. 0 //
30	7	19	28.43 43
40	7	19	29.25
50	7	19	30. 8
21. 0	7	19	30.51
10	7	19	31.34 //
20	7	19	32.16 43
30	7	19	32.59
40	7	19	33.41
50	7	19	34.24
22. 0	7	19	35. 6
10	7	19	35.49 //
20	7	19	36.31 43
30	7	19	37.14
40	7	19	37.56
50	7	19	38.39
23. 0	7	19	39.21
10	7	19	40. 3 //
20	7	19	40.46 42
30	7	19	41.28
40	7	19	42.10
50	7	19	42.52
24. 0	7	19	43.35
10	7	19	44.17 //
20	7	19	45. 0 42
30	7	19	45.42
40	7	19	46.24
50	7	19	47. 7

D. m.	S H	W H	Com. Pa. 1 11
25. 0	7	19	47.49
10	7	19	48.31 //
20	7	19	49.13 42
30	7	19	49.56
40	7	19	50.38
50	7	19	51.20
26. 0	7	19	52.01
10	7	19	52.43 //
20	7	19	53.25 42
30	7	19	54. 7
40	7	19	54.50
50	7	19	55.32
27. 0	7	19	56.13
10	7	19	56.55 //
20	7	19	57.37 42
30	7	19	58.20
40	7	19	59. 2
50	7	19	59.43
28. 0	8	20	0.15
10	8	20	1. 7 //
20	8	20	1.49 42
30	8	20	2.30
40	8	20	3.12
50	8	20	3.54
29. 0	8	20	4.36
10	8	20	5.18 //
20	8	20	6. 0 42
30	8	20	6.41
40	8	20	7.23
50	8	20	8. 4

D. m.	S. H	m H	Com. ps.
0.	0	8 30	8.46
10	8	20	9.28 "
20	8	20	10.10 42
30	8	20	10.52
40	8	20	11.34
50	8	20	12.16
1.	0	8 20	12.55
10	8	20	13.36 "
20	8	20	14.17 41
30	8	20	14.58
40	8	20	15.39
50	8	20	16.20
2.	0	8 20	17.4
10	8	20	17.45 "
20	8	20	18.26 41
30	8	20	19.07
40	8	20	19.48
50	8	20	20.29
3.	0	8 20	21.12
10	8	20	21.53 "
20	8	20	22.34 41
30	8	20	23.15
40	8	20	23.56
50	8	20	24.37
4.	0	8 20	25.20
10	8	20	26.01 "
20	8	20	26.42 41
30	8	20	27.23
40	8	20	28.04
50	8	20	28.45

D. m.	S. H	m H	Com. ps.
5.	0	8 20	29.27
10	8	20	30.08 "
20	8	20	30.49 41
30	8	20	31.30
40	8	20	32.11
50	8	20	32.52
6.	0	8 20	33.33
10	8	20	34.14 "
20	8	20	34.55 41
30	8	20	35.36
40	8	20	36.17
50	8	20	36.58
7.	0	8 20	37.38
10	8	20	38.19 "
20	8	20	39.00 41
30	8	20	39.41
40	8	20	40.22
50	8	20	41.03
8.	0	8 20	41.42
10	8	20	42.23 "
20	8	20	43.04 41
30	8	20	43.45
40	8	20	44.26
50	8	20	45.07
9.	0	8 20	45.46
10	8	20	46.26 "
20	8	20	47.06 40
30	8	20	47.46
40	8	20	48.26
50	8	20	49.06

D. m.	Ω H	ω H	Com. pts. H	
10.0	8	30	49.50	
10	8	20	50.30	"
20	8	20	51.10	
30	8	20	51.50	40
40	8	20	52.30	
50	8	20	53.10	
11.0	8	20	53.53	
10	8	20	54.33	"
20	8	20	55.13	40
30	8	20	55.53	
40	8	20	56.33	
50	8	20	57.13	
12.0	8	20	57.55	
10	8	20	58.35	"
20	8	20	59.15	40
30	8	20	59.55	
40	9	21	00.35	
50	9	21	1.15	
13.0	9	21	1.55	
10	9	21	2.35	"
20	9	21	3.15	40
30	9	21	3.55	
40	9	21	4.35	
50	9	21	5.15	
14.0	9	21	5.55	
10	9	21	6.35	"
20	9	21	7.15	40
30	9	21	7.55	
40	9	21	8.35	
50	9	21	9.15	

D. m.	Ω H	ω H	Com. pts. H	
15.0	9	21	9.55	
10	9	21	10.35	"
20	9	21	11.15	40
30	9	21	11.55	
40	9	21	12.35	
50	9	21	13.15	
16.0	9	21	13.53	
10	9	21	14.33	"
20	9	21	15.13	40
30	9	21	15.53	
40	9	21	16.33	
50	9	21	17.13	
17.0	9	21	17.51	
10	9	21	18.31	"
20	9	21	19.11	40
30	9	21	19.51	
40	9	21	20.31	
50	9	21	21.11	
18.0	9	21	21.48	
10	9	21	22.27	"
20	9	21	23.06	39
30	9	21	23.45	
40	9	21	24.24	
50	9	21	25.03	
19.0	9	21	25.45	
10	9	21	26.24	"
20	9	21	27.03	39
30	9	21	27.42	
40	9	21	28.21	
50	9	21	29.00	

D. m.	St H	HH	Com. Pts. H	
20.0	9	21	29.41	
10	9	21	30.26	"
20	9	21	30.59	39
30	9	21	31.88	
40	9	21	32.17	
50	9	21	32.56	
21.0	9	21	33.36	
10	9	21	34.15	"
20	9	21	34.54	39
30	9	21	35.33	
40	9	21	36.12	
50	9	21	36.51	
22.0	9	21	37.30	
10	9	21	38.9	"
20	9	21	38.48	39
30	9	21	39.27	
40	9	21	40.06	
50	9	21	40.45	
23.0	9	21	41.24	
10	9	21	42.03	"
20	9	21	42.42	39
30	9	21	43.21	
40	9	21	44.00	
50	9	21	44.39	
24.0	9	21	45.17	
10	9	21	45.56	"
20	9	21	46.35	39
30	9	21	47.14	
40	9	21	47.53	
50	9	21	48.32	

D. m.	St H	HH	Com. pts. H	
25.0	9	21	49.13	
10	9	21	49.50	"
20	9	21	50.29	39
30	9	21	51.08	
40	9	21	51.47	
50	9	21	52.26	
26.0	9	21	53.2	
10	9	21	53.41	"
20	9	21	54.19	39
30	9	21	54.58	
40	9	21	55.36	
50	9	21	56.14	
27.0	9	21	56.53	
10	9	21	57.32	"
20	9	21	58.10	39
30	9	21	58.49	
40	9	21	59.27	
50	10	22	00.06	
28.0	10	22	00.44	
10	10	22	01.22	"
20	10	22	02.00	38
30	10	22	02.39	
40	10	22	03.17	
50	10	22	03.55	
29.0	10	22	04.34	
10	10	22	05.12	"
20	10	22	05.50	38
30	10	22	06.29	
40	10	22	07.08	
50	10	22	07.46	

D. m.	H	H	Com. Pts.
0. 0	1022		8.24
10	1022		9.02 "
20	1022		9.40 38
30	1022		10.19
40	1022		10.57
50	1022		11.35
1. 0	1022		12.13
10	1022		12.51 "
20	1022		13.29 38
30	1022		14.07
40	1022		14.45
50	1022		15.23
2. 0	1022		16. 1
10	1022		16.39 "
20	1022		17.17 38
30	1022		17.55
40	1022		18.33
50	1022		19.11
3. 0	1022		19.49
10	1022		20.27 "
20	1022		21.05 38
30	1022		21.43
40	1022		22.21
50	1022		22.59
4. 0	1022		23.36
10	1022		24.14 "
20	1022		24.52 38
30	1022		25.30
40	1022		26.08
50	1022		26.46

D. m.	H	H	Com. Pts.
5. 0	1022		27.23
10	1022		28.01 "
20	1022		28.39 38
30	1022		29.16
40	1022		29.54
50	1022		30.32
6. 0	1022		31. 9
10	1022		31.46 "
20	1022		32.24 38
30	1022		33.01
40	1022		33.39
50	1022		34.17
7. 0	1022		34.55
10	1022		35.33 "
20	1022		36.10 38
30	1022		36.48
40	1022		37.25
50	1022		38.03
8. 0	1022		38.40
10	1022		39.18 "
20	1022		39.56 38
30	1022		40.33
40	1022		41.10
50	1022		41.48
9. 0	1022		42.25
10	1022		43.02 "
20	1022		43.40 38
30	1022		44.17
40	1022		44.54
50	1022		45.32

D. m.	H	H	Com. Pts.
10.0	1022		46.10
10	1022		46.47 "
20	1022		47.25 37
30	1022		48.03
40	1022		48.40
50	1022		49.17
11.0	1022		49.54
10	1022		50.31 "
20	1022		51.09 37
30	1022		51.46
40	1022		52.23
50	1022		53.01
12.0	1022		53.38
10	1022		54.15 "
20	1022		54.52 37
30	1022		55.29
40	1022		56.07
50	1022		56.44
13.0	1022		57.21
10	1022		57.58 "
20	1022		58.34 37
30	1022		59.12
40	1022		59.49
50	1123		00.27
14.0	1123		1.4
10	1123		1.41 "
20	1123		2.18 37
30	1123		2.55
40	1123		3.32
50	1123		4.09

D m.	H	H	Com. pts.
15.0	1123		4.47
10	1123		5.24 "
20	1123		6.01 37
30	1123		6.39
40	1123		7.16
50	1123		7.53
16.0	1123		8.29
10	1123		9.06 "
20	1123		9.43 37
30	1123		10.20
40	1123		10.57
50	1123		11.34
17.0	1123		12.11
10	1123		12.48 "
20	1123		13.25 37
30	1123		14.02
40	1123		14.39
50	1123		15.16
18.0	1123		15.53
10	1123		16.30 "
20	1123		17.07 37
30	1123		17.44
40	1123		18.21
50	1123		18.58
19.0	1123		19.34
10	1123		20.11 "
20	1123		20.48 37
30	1123		21.25
40	1123		22.02
50	1123		22.39

D. m.	度 H	分 H	Com. Pts. ' "
20.0	11	23	23.16
10	11	23	23.53 "
20	11	23	24.30 37
30	11	23	25.07
40	11	23	25.44
50	11	23	26.21
21.0	11	23	26.56
10	11	23	27.33 "
20	11	23	28.10 37
30	11	23	28.46
40	11	23	29.23
50	11	23	30.00
22.0	11	23	30.37
10	11	23	31.14 "
20	11	23	31.51 37
30	11	23	32.27
40	11	23	33.04
50	11	23	33.41
23.0	11	23	34.18
10	11	23	34.55 "
20	11	23	35.32 37
30	11	23	36.08
40	11	23	36.45
50	11	23	37.22
24.0	11	23	37.59
10	11	23	38.36 "
20	11	23	39.13 37
30	11	23	39.59
40	11	23	40.36
50	11	23	41.13

D. m.	度 H	分 H	Com. pts. ' "
25.0	11	23	41.40
10	11	23	42.17 "
20	11	23	42.54 37
30	11	23	43.30
40	11	23	44.07
50	11	23	44.43
26.0	11	23	45.20
10	11	23	45.57 "
20	11	23	46.34 37
30	11	23	47.10
40	11	23	47.47
50	11	23	48.23
27.0	11	23	49.00
10	11	23	49.37 "
20	11	23	50.14 37
30	11	23	50.50
40	11	23	51.27
50	11	23	52.03
28.0	11	23	52.40
10	11	23	53.17 "
20	11	23	53.54 37
30	11	23	54.30
40	11	23	55.06
50	11	23	55.43
29.0	11	23	56.20
10	11	23	56.57 "
20	11	23	57.34 37
30	11	23	58.10
40	11	23	58.46
50	11	23	59.23

A Catalogue of some Noted fix'd Stars, with their *Temporary Right Ascensions* to the Year 1686.

The first Column shews their Magnitudes. The second *Bayer's* Notes. The third the Numbers in *Ticho's* Catalogue. The fourth *Bayer's* Latin Names, and common English Names. The fifth their *Temporary Right Ascensions*.

Mag	Not.	Num	Stars Names.	Rt. Ascens. H. M. S.
	Bay.	Cat.		
		Tich		
3	ι	21	<i>In Extremitate Boreali Cauda Ceti</i> Northern in the Whales Tail	00.03.32
2	β	22	<i>In Extremitate Australi Cauda Ceti</i> Southern in the Whales Tail	00.28.17
3	α	16	<i>Præcedens in Eductione Cauda Ceti</i> Western in the Back of the Whale	00.54.13
3	θ	15	<i>Sequens in Eductione Cauda Ceti</i> Eastern in the Back of the Whale	01.08.40
3	ζ	14	<i>In medio corpore Ceti trium Septentrionalior</i> Northern in the Whales Belly	01.36.05
4	γ	1	<i>Prima Stella Arietis</i> First Star of Aries	01.36.22
4	β	2	<i>In sinistro Cornu Arietis</i> Second Star of Aries	01.37.04
3	α	19	<i>In Nodo Vinculorum Piscium</i> Br * in the Knot of the Fishes Line	01.45.49
3	α	14	<i>Lucida Arietis</i> Bright Star of Aries	01.49.29
3	δ	4	<i>Trium in Collo Ceti Australior</i> In the Whales Cheek	02.23.43
3	ε	10	<i>Lateris Sequentis Quadrilateri Boreæ in Pectore Ceti</i> On the Whales Brest	02.24.40
3	γ	3	<i>Media Trium in Collo Ceti</i> In the Whales Mouth.	02.27.16

2	•	2	<i>Lucida in Naribus, Menkar, Ceti</i> Whales Jaw	h 1 11 02.45.50
3	•	32	<i>Lucida Pleiadum Tauri</i> Brightest of the Pleiades	03.29.01
3	γ	11	<i>In Vultu Tauri illa in Naribus</i> The lowest of the Hyades	04.01.50
3	•	15	<i>Ad Oculum Boreum Tauri</i> North Eye of Taurus	04.10.08
1	•	14	<i>Tauri Oculus Austrinus. Aldebaran</i> Bulls South Eye	04.17.56
3	β	2	<i>Supra Pedem Orionis in Flumine Erid.</i> Orions Foot in Eridanus	04.52.32
1	β	35	<i>Pes Dexter Orionis. Rigel</i> Bright Star in Orions Foot	04.59.29
2	•		<i>Tauri Cornu Septentrionale</i> Bulls North Horn	05.06.27
2	γ	3	<i>In humero Dextro Orionis Duarum Borealior</i> First or left Shoulder of Orion	05.08.22
2	•	26	<i>In Baltheo Orionis Fulgentium 3um precedens.</i> First in Orions Belt	05.15.55
3	ζ	19	<i>Extrema Cornu Australis Tauri.</i> Bulls South Horn	05.18.50
2	•	27	<i>Media Balthei Orionis</i> The second or middle in Orions Belt	05.20.08
2	ζ	28	<i>Sequens seu tertia Balthei Orionis.</i> Third or last in Orions Belt	05.24.30
1	•	2	<i>Lucida in Humero sinistro Orionis</i> The latter or right Shoulder of Orion	05.38.07
3	•	15	<i>Calx Castoris</i> The Heel of Gemini	05.56.05

2	♂	9	<i>In Collario Canis Majoris</i> The fore-foot of the Great Dog	06.09.13
2	γ	17	<i>In summo Pede sinistro sequentis</i> II Bright Star on the Foot of II	06.19.25
1	α	1	<i>Syrus seu Canis Major</i> The Great Dog Star	06.31.11
3	ζ	11	<i>In Dextro-Genu Herculis</i> II Left Knee of Gemini	06.45.38
3	γ	4	<i>Ad Aurem Dextram Canis Majoris</i> On the Neck of Canis Major	07.11.00
1	α	2	<i>Procyon seu Canis Minor</i> The Little Dog	07.22.53
2	α	12	<i>Cōr Hydrae, Alphard</i> Hydra's Heart	09.12.07
3	γ	7	<i>Trium in Cervice Leonis Australior</i> Southern in the Neck of Leo	09.50.22
1	α	8	<i>Lor Leonis</i> Lions Heart	09.51.31
3	ζ	5	<i>Trium in Cervice Leonis Borealiōr</i> Northern in the Neck of Leo	09.59.06
2	γ	6	<i>Trium in Cervice Leonis Media. Juba</i> Lions Crest	10.02.27
2	δ	20	<i>In Lumbis Leonis Duarum quæ Sequitur</i> In the Loyn of Leo	10.57.21
3	θ	21	<i>In Vertebra Femoris Dextri Leonis</i> On the Buttock of Leo	10.57.50
1	θ	27	<i>Cauda Leonis</i> Lions Tail	11.32.58
3	ι	23	<i>In Femore Dextro Leonis</i> On the Thigh of Leo	11.07.44

3	β	5	<i>In Extremo Ala sinistra Virginis</i> Above the left wing of Virgo	h 1 11 11.34.31
3	γ	7	<i>Sub strophio seu Castula Virginis</i> On the left wing of Virgo	12.25.43
3	δ	10	<i>Ad Cingulum Virginis</i> In the Girdle of Virgo	12.39.58
3	ϵ	13	<i>Ala Dextra \propto precedens. Vindemiatrix</i> Virgo's right wing. Vindemiatrix	12.46.46
1	α	14	<i>Spica Virginis</i> Virgins Spike	13.08.46
3	ζ	15	<i>Sub Perizomate ad pedem dextrum \propto</i> On the Buttock of Virgo	13.13.38
3	κ	20	<i>In sinistra Tibia Bootis trium Borealis</i> Most Northern on the left knee of Bootes	13.32.33
1	α	23	<i>Arcturus</i> Arcturus	14.01.24
3	ζ	19	<i>In Dextra Tibia Bootis</i> Bend of Bootes right knee	14.26.20
2	α	1	<i>Lanx Australis</i> Southern Ballance	14.33.46
2	β	3	<i>Lanx Borealis</i> Northern Ballance	15.00.15
3	γ	7	<i>In media lance Boreali Prima & Superior</i> Middle of the Northern Ballance	15.18.13
2	α	9	<i>Precedens trium in medio neru colli serpentis</i> Bright Star in the Serpents neck	15.28.54
2	β	1	<i>In fronte \propto ad Boream prima fulgens</i> Highest in Scorpio's forehead	15.46.04
3	δ	2	<i>In dextra manu serpentarii Borealis</i> Northern in the left hand of Serpentarius	15.58.23

3	•	8	<i>Australior in Dextra manu Serpentarii</i> Southern in the left hand of Serpentarius	h 1 11 16.01.57
3	7	3	<i>In sinist. Humero Herculis trium Australior</i> Hercules Right Shoulder	16.08.07
1	•	8	<i>Cor Scorpii</i> Scorpius Heart	16.10.18
3	8	2	<i>In sinist. Humero Herculis trium Boreali</i> Northern in Hercules Right Shoulder	16.16.41
3	2	19	<i>In Femore Dextro Serpentarii</i> Serpentarius left knee	16.19.59
3	0	13	<i>In Femore sinistro Serpentarii 3^{um} Boreali</i> Right Ham of Serpentarius	16.52.02
3	•	1	<i>Caput Herculis</i> Hercules Head	17.00.23
3	8	5	<i>Herculis Humerus Dexter</i> Left Shoulder of Hercules	17.01.57
2	•	1	<i>Caput Serpentarii</i> Serpentarius Head	17.20.22
3	8	2	<i>In sinist. Humero Serpentarii 2^{um} Superior</i> Northern in Serpentarius right Shoulder	17.27.55
3	7	3	<i>In eodem duarum Inferior</i> Southern in Serpentarius Right Shoulder	17.32.23
3	2	16	<i>Prima post Serpentarium in Cauda Serpentis</i> Last but two in the Serpents Tail	17.44.09
3	•	17	<i>In Penultima Flexione Cauda Serpentis</i> Last but one in the Serpents Tail	18.05.20
3	0	18	<i>Extrema Cauda Serpentis</i> Last in the Serpents Tail	18.40.37
3	•	11	<i>Extrema Cauda Aquile</i> Last in the Eagles Tail	18.51.06

3	γ	5	<i>Ad Originem Alae Dextrae Aquila</i> Pinion of Aquila's Right Wing	h 1 11 19.31.38.
2	α	3	<i>Lucida in Eductione Colli Aquila</i> Aquila's Bright Star	19.35.51
3	β	2	<i>In Collo Aquila</i> Neck of Aquila	19.39.57
3	α	1	<i>Septentrionalis duplex in Cornu Capricorni</i> Northern in the Horn of Capricorn	20.00.50
3	α	1	<i>Lucida Caudae Delphini trium Antecedens.</i> Dolphins Tail	20.18.23
3	β	4	<i>In Humero Dextro Aquarii</i> Left Shoulder of Aquary	21.14.58
3	γ	23	<i>In Eductione Caudae Capricorni Antecedens</i> First in the Tail of Capricorn	21.21.57
3	α	1	<i>In Rictu Pegasi. Emis.</i> Mouth of Pegasus	21.28.44
3	β	24	<i>In Eductione Caudae Capricorni Succedens</i> Last in the Tail of Capricorn	21.30.07
3	α	2	<i>Humerus sinister Aquarii</i> Right Shoulder of Aquary	21.49.43
3	γ	9	<i>In Lavo Cubito Aquarii</i> Right Elbow of Aquary	22.05.37
3	β	18	<i>In sinistra Tibia Aquarii</i> Right shin or knee of Aquary	22.38.15
2	β	18	<i>In Eductione Cruris Pegasi. Scheat</i> Pegasus Thigh	22.48.43
2	α	17	<i>Ala Pegasi. Marchab</i> Pegasus Wing	22.49.05
2	γ	19	<i>In Exteema Ala Pegasi</i> End of Pegasus Wing	23.57.03

TABLES
OF
ÆQUATION
OF
Natural Days.

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TABLES
OF
EQUATION
OF

Natural Days

A Table of Equation of Natural Days.

Day	Jan.	Diff.	Feb.	March	Apr.	May	June
1	8.53	22	14.54	10.16	18.0	5.16	4.12
2	9.17	23	14.53	9.58	18.0	3.35	4.14
3	9.40	20	14.51	9.41	17.0	1.18	4.16
4	10.02	21	14.48	9.24	17.0	0.01	4.17
5	10.23	21	14.45	9.06	18.0	1.14	4.16
6	10.44	18	14.40	8.49	0.29	4.15	0.07
7	11.02	18	14.35	8.32	0.43	4.14	0.06
8	11.20	18	14.29	8.14	0.58	4.12	0.19
9	11.38	17	14.23	7.55	1.12	4.10	0.32
10	11.55	17	14.16	7.37	1.26	4.08	0.44
11	12.12	16	14.07	7.18	1.40	4.05	0.57
12	12.28	15	13.59	6.58	1.52	4.01	1.10
13	12.43	15	13.50	6.39	2.03	3.57	1.23
14	12.58	13	13.41	6.20	2.15	3.52	1.36
15	13.11	12	13.31	6.01	2.26	3.47	1.50
16	13.23	12	13.20	5.42	2.37	3.42	2.02
17	13.35	11	13.09	5.23	2.48	3.35	2.14
18	13.46	11	12.57	5.04	2.57	3.28	2.26
19	13.57	9	12.45	4.46	3.06	3.21	2.38
20	14.06	9	12.31	4.27	3.15	3.13	2.50
21	14.15	8	12.17	4.08	3.23	3.04	3.02
22	14.23	6	12.04	3.50	3.30	2.56	3.14
23	14.29	6	11.56	3.31	3.37	2.47	3.25
24	14.35	5	11.37	3.12	3.44	2.37	3.35
25	14.40	4	11.22	2.53	3.50	2.28	3.45
26	14.44	4	11.06	2.35	3.55	2.18	3.57
27	14.48	2	10.50	2.16	3.59	2.07	4.07
28	14.50	1	10.34	1.58	3.53	1.56	4.17
29	14.51	1	10.17	1.41	3.47	1.45	4.25
30	14.52	1	10.01	1.24	3.40	1.34	4.33
31	14.53	1	9.45	1.07	3.32	1.22	

A Table of Equation of Natural Days.

Day	July.	X	August	X	Sept.	X	Octob.	X	Novemb	X	Decemb	X
	I	II	III	I	II	III	I	II	III	I	II	III
1	4. 43	8	4. 36	9	3. 41	21	13. 13	13	15. 28	8	5. 51	28
2	4. 51	7	4. 27	9	4. 02	21	13. 26	14	15. 20	9	5. 23	28
3	4. 58	7	4. 18	12	4. 23	21	13. 40	13	15. 11	10	4. 55	29
4	5. 5	7	4. 06	11	4. 44	20	13. 53	13	15. 01	10	4. 26	30
5	5. 12	6	3. 55	11	5. 04	20	14. 06	13	14. 51	12	Subtract 56	29
6	5. 18	6	3. 44	13	5. 24	21	14. 19	12	14. 39	13	Subtract 27	29
7	5. 24	5	3. 31	13	5. 45	20	14. 31	11	14. 26	13	3. 58	30
8	5. 29	4	3. 18	13	6. 05	21	14. 42	10	14. 13	13	2. 28	30
9	5. 33	4	3. 05	13	6. 26	20	14. 52	09	14. 00	13	2. 58	30
10	5. 37	4	2. 52	14	6. 46	21	15. 01	9	13. 45	15	1. 28	30
11	5. 41	3	2. 38	14	7. 07	21	15. 10	8	13. 29	16	0. 58	31
12	5. 44	2	2. 24	16	7. 28	21	15. 18	8	13. 13	17	0. 28	30
13	5. 46	2	2. 08	16	7. 49	20	15. 26	8	12. 56	18	0. 03	30
14	5. 48	1	1. 52	16	8. 09	22	15. 34	6	12. 38	15	0. 33	30
15	5. 49	0	1. 36	16	Subtract 30	20	Subtract 40	6	Subtract 19	15	0. 03	30
16	5. 49	0	1. 20	16	8. 50	19	Subtract 46	5	Subtract 00	21	1. 33	29
17	5. 49	1	1. 04	17	9. 09	20	15. 51	4	11. 39	21	2. 02	30
18	5. 48	1	0. 47	17	9. 29	19	15. 55	3	11. 18	22	2. 32	29
19	5. 47	2	0. 30	18	9. 48	18	15. 58	2	10. 56	22	3. 01	30
20	5. 45	2	0. 12	17	10. 6	19	16. 00	2	10. 34	23	3. 31	29
21	5. 43	3	0. 05	19	10. 25	19	16. 02	1	10. 11	24	+ Add 00	28
22	5. 40	4	0. 24	19	10. 44	19	16. 03	1	9. 47	24	+ Add 28	28
23	5. 36	4	0. 43	20	11. 03	17	16. 04	0	9. 23	24	4. 56	27
24	5. 32	4	1. 03	19	11. 20	17	16. 04	2	8. 59	25	5. 23	27
25	5. 28	6	Subtract 22	19	11. 37	17	16. 02	3	8. 34	26	5. 50	27
26	5. 22	7	1. 41	19	11. 54	17	15. 59	3	8. 08	27	6. 17	27
27	5. 15	7	2. 00	20	12. 11	16	15. 56	4	7. 41	26	6. 44	26
28	5. 08	8	2. 20	20	12. 27	16	15. 52	5	7. 15	27	7. 10	26
29	5. 00	7	2. 40	20	12. 43	16	15. 47	6	6. 48	29	7. 36	24
30	4. 53	8	3. 00	20	12. 59	15	15. 41	7	6. 19	8	8. 00	24
31	4. 45	3	3. 20				15. 34				8. 24	

CALCULATION
OF
Hours and Minutes for an
HORIZONTAL-DIAL
DUBLIN Lat. $53^{\circ}. 20'$.

COLLECTION

OF

HORISONTAL

FOR

Calculation of Hours and Minutes for an Horizontal Dial.

Dublin Lat: $53^{\circ} 26'$.

Time from Noon	Hour dist. on the Plain	Time from Noon	Hour dist. on the Plain	Time from Noon	Hour dist. on the Plain	Time from Noon	Hour dist. on the Plain
h. m.	h. m.	h. m.	h. m.	h. m.	h. m.	h. m.	h. m.
XII. 00 00 . 00		XII. 30 06 . 11		I. 00 12 . 8		I. 30 18 . 23	
1 00 12 . 12		1 30 14 . 14		1 12 . 20		31 18 . 35	
2 00 24 . 24		2 30 26 . 26		2 12 . 33		32 18 . 48	
3 00 36 . 36		3 30 38 . 38		3 12 . 45		33 19 . 1	
4 00 48 . 48		4 30 50 . 50		4 12 . 57		34 19 . 14	
5 01 00 . 00		5 30 2 . 2		5 13 . 9		35 19 . 27	
6 01 12 . 12		6 30 14 . 14		6 13 . 22		36 19 . 39	
7 01 24 . 24		7 30 27 . 27		7 13 . 34		37 19 . 52	
8 01 36 . 36		8 30 39 . 39		8 13 . 47		38 20 . 4	
9 01 48 . 48		9 30 51 . 51		9 13 . 59		39 20 . 17	
10 02 00 . 00		10 30 3 . 3		10 14 . 11		40 20 . 30	
11 02 12 . 12		11 30 15 . 15		11 14 . 24		41 20 . 43	
12 02 25 . 25		12 30 28 . 28		12 14 . 37		42 20 . 56	
1 02 37 . 37		1 30 39 . 39		13 14 . 49		43 21 . 9	
2 02 49 . 49		2 30 51 . 51		14 14 . 1		44 21 . 22	
3 02 1 . 1		3 30 4 . 4		15 15 . 14		45 21 . 35	
4 02 13 . 13		4 30 16 . 16		16 15 . 26		46 21 . 48	
5 02 25 . 25		5 30 28 . 28		17 15 . 39		47 22 . 1	
6 02 37 . 37		6 30 40 . 40		18 15 . 51		48 22 . 14	
7 02 49 . 49		7 30 52 . 52		19 16 . 4		49 22 . 27	
8 02 1 . 1		8 30 5 . 5		20 16 . 17		50 22 . 40	
9 02 13 . 13		9 30 17 . 17		21 16 . 29		41 22 . 53	
10 02 25 . 25		10 30 29 . 29		22 16 . 41		52 23 . 6	
11 02 37 . 37		11 30 42 . 42		23 16 . 54		53 23 . 19	
12 02 49 . 49		12 30 54 . 54		24 17 . 7		54 23 . 32	
1 02 59 . 1		1 30 11 . 6		25 17 . 20		55 23 . 45	
2 03 13 . 13		2 30 18 . 18		26 17 . 32		56 23 . 58	
3 03 25 . 25		3 30 31 . 31		27 17 . 45		57 24 . 11	
4 03 37 . 37		4 30 43 . 43		28 17 . 57		58 24 . 24	
5 03 49 . 49		5 30 55 . 55		29 18 . 10		59 24 . 38	
6 03 1 . 1		6 30 12 . 8		30 18 . 23		60 24 . 51	

Calculation of Hours and Minutes for an Horizontal Dial:

Dublin Lat: $53^{\circ} 20'$.

Time from Noon	Hour diff. on the Plain	Time from Noon	Hour diff. on the Plain	Time from Noon	Hour diff. on the Plain	Time from Noon	Hour diff. on the Plain
h' o'		h' o'		h' o'		h' o'	
II. 00	24 . 51	II. 30	31 . 37	III. 00	38 . 44	III. 30	46 . 16
1	25 . 5	31	31 . 51	1	38 . 59	31	46 . 32
2	25 . 18	32	32 . 4	2	39 . 14	32	46 . 47
3	25 . 31	33	32 . 18	3	39 . 28	33	47 . 3
4	25 . 44	34	32 . 32	4	39 . 43	34	47 . 19
5	25 . 57	35	32 . 46	5	39 . 57	35	47 . 35
6	26 . 10	36	33 . 0	6	40 . 12	36	47 . 50
7	26 . 24	37	33 . 14	7	40 . 27	37	48 . 9
8	26 . 38	38	33 . 28	8	40 . 42	38	48 . 21
9	26 . 51	39	33 . 43	9	40 . 57	39	48 . 37
10	27 . 4	40	33 . 57	10	41 . 12	40	48 . 53
11	27 . 18	41	34 . 11	11	41 . 27	41	49 . 9
12	27 . 31	42	34 . 25	12	41 . 42	42	49 . 25
13	27 . 44	43	34 . 39	13	41 . 57	43	49 . 40
14	27 . 58	44	34 . 53	14	42 . 12	44	49 . 56
15	28 . 11	45	35 . 8	15	42 . 27	45	50 . 12
16	28 . 25	46	35 . 22	16	42 . 42	46	50 . 28
17	28 . 39	47	35 . 36	17	42 . 57	47	50 . 44
18	28 . 52	48	35 . 50	18	43 . 12	48	51 . 0
19	29 . 6	49	36 . 5	19	43 . 27	49	51 . 16
20	29 . 19	50	36 . 19	20	43 . 42	50	51 . 33
21	29 . 33	51	36 . 34	21	43 . 57	51	51 . 49
22	29 . 47	52	36 . 48	22	44 . 13	52	52 . 5
23	30 . 0	53	37 . 2	23	44 . 29	53	52 . 21
24	30 . 14	54	37 . 17	24	44 . 44	54	52 . 37
25	30 . 28	55	37 . 31	25	44 . 59	55	52 . 54
26	30 . 41	56	37 . 46	26	45 . 15	56	53 . 10
27	30 . 55	57	38 . 0	27	45 . 30	57	53 . 26
28	31 . 9	58	38 . 15	28	45 . 45	58	53 . 43
29	31 . 23	59	38 . 29	29	46 . 1	59	53 . 59
30	31 . 37	60	38 . 44	30	46 . 16	60	54 . 15

Calculation of Hours and Minutes for an Horizontal Dial:

Dublin Lat: 53° 20'.

Time from Noon	Hour diff. on the Plain	Time from Noon	Hour diff. on the Plain	Time from Noon	Hour diff. on the Plain	Time from Noon	Hour diff. on the Plain
h ' o '		h ' o '		h ' o '		h ' o '	
III. 00	54 . 15	III. 30	62 . 42	V. 00	71 . 32	V. 30	80 . 40
1	54 . 32	31	62 . 59	1	71 . 50	31	80 . 59
2	54 . 48	32	63 . 16	2	72 . 8	32	81 . 18
3	55 . 5	33	63 . 33	3	72 . 26	33	81 . 37
4	55 . 21	34	63 . 51	4	72 . 44	34	81 . 55
5	55 . 37	35	64 . 8	5	73 . 2	35	82 . 14
6	55 . 54	36	64 . 26	6	73 . 20	36	82 . 32
7	56 . 11	37	64 . 43	7	73 . 38	37	82 . 51
8	56 . 28	38	65 . 1	8	73 . 56	38	83 . 9
9	56 . 45	39	65 . 18	9	74 . 14	39	83 . 28
10	57 . 1	40	65 . 36	10	74 . 33	40	83 . 47
11	57 . 18	41	65 . 53	11	74 . 51	41	84 . 5
12	57 . 35	42	66 . 11	12	75 . 9	42	84 . 24
13	57 . 51	43	66 . 28	13	75 . 28	43	84 . 43
14	58 . 8	44	66 . 46	14	75 . 46	44	85 . 2
15	58 . 25	45	67 . 4	15	76 . 4	45	85 . 20
16	58 . 42	46	67 . 22	16	76 . 23	46	85 . 38
17	58 . 59	47	67 . 39	17	76 . 41	47	85 . 57
18	59 . 16	48	67 . 57	18	76 . 59	48	86 . 16
19	59 . 32	49	68 . 14	19	77 . 17	49	86 . 34
20	59 . 49	50	68 . 32	20	77 . 36	50	86 . 53
21	60 . 7	51	68 . 50	21	77 . 55	51	87 . 11
22	60 . 24	52	69 . 8	22	78 . 13	52	87 . 31
23	60 . 41	53	69 . 26	23	78 . 31	53	87 . 50
24	60 . 58	54	69 . 44	24	78 . 50	54	88 . 9
25	61 . 15	55	70 . 2	25	79 . 8	55	88 . 28
26	61 . 32	56	70 . 20	26	79 . 27	56	88 . 47
27	61 . 50	57	70 . 38	27	79 . 45	57	89 . 6
28	62 . 7	58	70 . 56	28	80 . 4	58	89 . 25
29	62 . 24	59	71 . 14	29	80 . 22	59	89 . 44
30	62 . 42	60	71 . 32	30	80 . 40	60	90 . 00

N

A Table of the most noted *Circumpolar Stars*, or those that set not in the *Lat.* 53°. 20'. for shewing the time of Night by a Line and Plummert, to the year 1680, and serves for 20 years to come.

Mag	Stars Names.	Rt. Ascension in Time under the Pole star	Diff. in Time betwixt the pole star & pole star	Azimuth un- der the Pole star	Coast
		H ' "	H ' "	o ' "	
	Urfa Major				
2	<i>On his Back</i>	00.03.53	0 .04.26	0 .37.12	E
3	<i>Penult in Dracos Tail</i>	00.20.17	0 .03.19	0 .19.12	E
2	<i>Cor Caroli</i>	00.36.44	0 .0 .05	0 .00.54	E
2	<i>His Rump Aliot</i>	00.39.25	0 .0 .14	0 .02.06	W
2	<i>Middle of its Tail</i>	01.07.02	0 .03.44	0 .33.00	W
2	<i>End of the Tail</i>	01.28.56	0 .05.49	0 .57.00	W
2	<i>Last but two in Dracos Tail</i>	01.47.40	0 .10.58	1 .16.48	W
	Urfa Minor				
2	<i>Higher Guard</i>	02.26.05	0 .24.10	1 .55.48	W
3	<i>Lower Guard</i>	02.57.15	0 .26.50	2 .24.36	W
	Draco				
3	<i>27 in Ticho's Cat.</i>	02.59.49	0 .18.01	2 .27.00	W
3	<i>26</i>	03.35.17	0 .21.04	2 .55.48	W
3	<i>25</i>	03.54.54	0 .24.39	3 .10.12	W
3	<i>20</i>	03.58.34	0 .30.43	3 .12.36	W
3	<i>24</i>	04.36.43	0 .30.49	3 .34.48	W
3	<i>First in the Head</i>	04.59.30	0 .23.42	3 .45.00	W
2	<i>Bright in the Head</i>	05.25.01	0 .24.11	3 .53.24	W
1	<i>Lutida Lyra</i>	06.05.34	0 .20.30	4 .00.36	W

		H	I	"	H	I	"	0	I	"	
3	N. in the upper turn	06.37.04	0		.35.44	4		20.01.12			W
3	S. in the upper turn	07.11.22	0		.37.53	3		56.24			W
	Cygnus										
3	Upper Wing	07.13.26	0		.21.49	3		56.24			W
3	Her Brest	07.51.26	0		.19.20	3		45.00			W
4	N. Cepheus R. Arm	07.55.07	0		.28.46	3		44.00			W
2	Swans Tail	08.10.30	0		.19.59	3		37.12			W
	Cepheus										
4	S. on his R. Arm	08.11.32	0		.26.57	3		36.36			W
2	Right-shoulder	08.45.35	0		.25.02	3		18.36			W
3	His Girdle	08.53.30	0		.30.44	3		13.48			W
4	N. in his Cap	09.41.52	0		.18.16	2		40.48			W
4	Before his Head	10.09.17	0		.15.56	2		19.12			W
4	Left shoulder	10.20.44	0		.17.48	2		09.36			W
3	His Knee	11.10.00	0		.17.59	1		25.48			W
	Calliopea										
3	Her Chair	11.46.07	0		.05.54	0		51.00			W
4	Seat of the Chair	12.11.48	0		.03.14	0		25.48			W
4	Head	12.17.20	0		.02.07	0		20.24			W
3	Brest	12.20.48	0		.01.50	0		16.48			W
4	Girdle	12.28.46	0		.00.59	0		09.00			W
3	Hipp	12.37.50	0		.00.02	0		00.10			W
3	Knee	13.08.45	0		.03.41	0		31.12			E
3	Legg	13.39.38	0		.07.53	1		01.12			E
2	Foot of Androm.	13.50.37	0		.06.19	1		12.00			E

Perseus		H 1 "	H 1 "	O 1 "	
4	Left Shoulder	14.33.32	O .10.52	1 .51.36	E
3	Bend of his side	14.43.40	O .11.03	2 .00.36	E
3	Right Shoulder	14.55.31	O .13.42	2 .10.48	E
3	Medusa's Head	14.59.14	O .11.45	2 .13.48	E
2	Bright * in his side	15.14.59	O .14.29	2 .26.24	E
4	Left Thigh	15.37.56	O .14.35	2 .44.24	E
3	Left Knee	15.51.39	O .14.57	2 .54.36	E
4	Before his Right Knee	16.01.08	O .18.05	3 .00.36	E
4	Middle of his Right Knee	16.09.47	O .18.06	3 .06.36	E
Auriga					
4	Left Elbow	16.58.15	O .19.27	3 .33.36	E
1	Capella	17.13.55	O .20.49	3 .40.48	E
2	Right Shoulder	17.58.07	O .21.55	3 .55.12	E
Ursa Major					
4	His Lipp	20.31.52	O .27.38	3 .37.48	E
3	N. in his Right Foot	20.50.49	O .20.53	3 .28.12	E
3	S. in the same	20.55.23	O .20.13	3 .21.00	E
3	Left Knee	21.24.15	O .19.57	3 .07.12	E
4	Higher in his hind Foot	22.11.32	O .13.56	2 .30.36	E
2	Lower Leader in □	22.54.50	O .12.57	1 .50.24	E
2	Higher Leader	22.57.59	O .14.20	1 .47.24	E
3	Last in Draco's Tail	23.24.13	O .13.36	1 .20.24	E
2	Bears Left Thigh	23.43.36	O .06.38	0 .59.24	E

F I N I S.

